# **The Matérn Model: A Journey through Statistics, Numerical Analysis and Machine Learning**

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*Abstract.* The Matérn model has been a cornerstone of spatial statistics for more than half a century. More recently, the Matérn model has been exploited in disciplines as diverse as numerical analysis, approximation theory, computational statistics, machine learning, and probability theory. In this article we take a Matérn-based journey across these disciplines. First, we reflect on the importance of the Matérn model for estimation and prediction in spatial statistics, establishing also connections to other disciplines in which the Matérn model has been influential. Then, we position the Matérn model within the literature on big data and scalable computation: the SPDE approach, the Vecchia likelihood approximation, and recent applications in Bayesian computation are all discussed. Finally, we review recent devlopments, including flexible alternatives to the Matérn model, whose performance we compare in terms of estimation, prediction, screening effect, computation, and Sobolev regularity properties.

*Keywords*: Approximation Theory, Compact Support, Covariance, Kernel, Kriging, Machine Learning, Maximum Likelihood, Reproducing Kernel Hilbert Spaces, Spatial Statistics, Sobolev Spaces.

# **1. INTRODUCTION**

 This paper serves two purposes: On the one hand, we provide a panoramic view, across several disciplines, of the Matérn model. On the other hand, the paper illustrates the role of the Matérn model in several disciplines, while discussing alternative or more general models and their relevance to many aspects of statistical modeling, estima-tion, prediction, computational statistics, numerical anal-

<sup>8</sup> ysis, and machine learning.

 A historical account of the Matérn model is provided 10 by Guttorp and Gneiting  $[69]$ . The Matérn model – also called the Matérn *covariance function*, or the Matérn *ker- nel*, depending on context – is commonly attributed to Matérn  $[109]$ , but can be found under alternative names in different branches of the scientific literature. The use of the Matérn model is widespread, and it is impossible to cover all its diverse applications here; our review focuses on a selection of applications that are of especial interest and significance. Specifically, we aim to cover

- <sup>19</sup> 1. estimation and prediction using the Matérn model <sup>20</sup> in statistics, with emphasis on maximum likelihood <sup>21</sup> estimation, Kriging prediction, and the associated <sup>22</sup> screening effect;
- <sup>23</sup> 2. applications of the Matérn model in
- <sup>24</sup> a) computational statistics, including the stochas-<sup>25</sup> tic differential equation (SDE) and stochas-<sup>26</sup> tic partial differential equation (SPDE) ap-<sup>27</sup> proaches, likelihood approximation, inference <sup>28</sup> of partial differential equations (PDEs) and <sup>29</sup> Charles Stein's method;
- <sup>30</sup> b) statistical modeling, including non-standard <sup>31</sup> scenarios, for instance when isotropy and sta-

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 This article is novel, in being the first to take a broad view of the scientific literature through the lens of the Matérn model. In particular, we do *not* attempt a review of covari- ance functions in general. Recent reviews provide a quite exhaustive panorama of covariance models, from space to space-time [\[128\]](#page-21-0), to multivariate covariance functions  $62 \quad [58]$  $62 \quad [58]$ , and covariance-based modeling on spheres and man- ifolds [\[123\]](#page-20-1). In addition, while there are many fascinat- ing applications of the Matérn model across the scientific landscape, we cannot hope to do justice to them all. Our emphasis is therefore limited to methodological and the- oretical issues which we hope are of relevance across a wide range of disciplines in which the Matérn model is <sup>69</sup> used.

# <span id="page-1-1"></span><sup>70</sup> **1.1 Setting and Notation**

<sup>73</sup> definite.

Throughout, bold letters refer to vectors and matrices, and the transpose operator is denoted ⊤. Let  $d \in \mathbb{N}$  and let  $Z = \{Z(\boldsymbol{x}), \ \boldsymbol{x} \in \mathbb{R}^d\}$  be a real-valued Gaussian random field, having zero mean and and *covariance function* K :  $\mathbb{R}^d \times \mathbb{R}^d \to \mathbb{R}$  defined via  $K(\boldsymbol{x}, \boldsymbol{y}) := \text{Cov}(Z(\boldsymbol{x}), Z(\boldsymbol{y})).$ Covariance functions are symmetric and positive definite, where in this paper the term *positive definite* is understood as

(1) 
$$
\sum_{i=1}^{n} \sum_{j=1}^{n} c_i K(\boldsymbol{x}_i, \boldsymbol{x}_j) c_j \ge 0
$$

<sup>71</sup> for all  $c_i \in \mathbb{R}$ , all  $n \in \mathbb{N}$  and all  $x_i \in \mathbb{R}^d$ . If the inequal- $72$  ity above is strict, then K will be called strictly positive

Each symmetric positive definite function  $K : \mathbb{R}^d \times$  $\mathbb{R}^d \to \mathbb{R}$  defines *translate* functions  $K(\boldsymbol{x}, \cdot)$  on  $\mathbb{R}^d$ , for all  $x \in \mathbb{R}^d$ . In addition, one can define an inner product on two translates by

<span id="page-1-0"></span>
$$
(2) \quad \langle K(\boldsymbol{x},\cdot),K(\boldsymbol{y},\cdot)\rangle_{\mathcal{H}(\mathcal{K})}:=K(\boldsymbol{x},\boldsymbol{y}),\ \boldsymbol{x},\,\boldsymbol{y}\in\mathbb{R}^d,
$$

in terms of  $K$  itself. This extends to all linear combinations of translates and *generates*, by completion, a Hilbert space  $\mathcal{H}(\mathcal{K})$  of functions on  $\mathbb{R}^d$ . This space is called the *native* space for K. Notice that the Hilbert space allows for continuous point evaluations  $\delta_{\mathbf{x}} : f \mapsto f(\mathbf{x})$  via a *reproduction formula*

<span id="page-1-2"></span>(3) 
$$
f(\mathbf{x}) = \langle f, K(\mathbf{x}, \cdot) \rangle_{\mathcal{H}(K)}, \mathbf{x} \in \mathbb{R}^d, f \in \mathcal{H}(K)
$$

that follows from  $(2)$ . Then  $\mathcal{H}(K)$  is called a *reproducing kernel Hilbert space (RKHS)* with *kernel* K. In particular, the translates  $K(\mathbf{x}, \cdot)$  lie in  $\mathcal{H}(K)$ , forming its completion and being the Riesz representers of delta functionals  $\delta_x$ . They are central to machine learning, numerical analysis and approximation theory, since [\(2\)](#page-1-0) allows inner products in the abstract space  $\mathcal{H}(K)$  to be explicitly computable using the kernel - the so-called *kernel trick*. See Section [6.1](#page-9-0) and [\[167\]](#page-21-1) for more detail. For a positive definite and stationary kernel  $K$ , its Fourier transform  $K$  can be used to recast the inner product [\(2\)](#page-1-0) on the Hilbert space  $\mathcal{H}(K)$ by

<span id="page-1-3"></span>(4) 
$$
\langle f, g \rangle_{\mathcal{H}(K)} = \int_{\mathbb{R}^d} \frac{\hat{f}(\omega) \overline{\hat{g}(\omega)}}{\hat{K}(\omega)} d\omega, \ f, g \in \mathcal{H}(K),
$$

 $74$  up to a constant factor. Here,  $\overline{g}$  denotes the complex con- $75$  jugate of a function g, and  $\hat{g}$  its Fourier transform. Note how the spectrum of  $K$  penalizes the spectrum of the functions in  $\mathcal{H}(K)$ . Roughly, the Hilbert space  $\mathcal{H}(K)$ <sup>78</sup> consists of functions f for which  $\hat{f}/\sqrt{\hat{K}}$  is square in- $\tau$ <sup>9</sup> tegrable over  $\mathbb{R}^d$ . The subtle connections of the Hilbert so space  $\mathcal{H}(K)$  to sample paths of Gaussian processes with  $81$  covariance function K will come up at many places in  $\alpha$  this paper, e.g. in Sections [2,](#page-2-0) [4.4,](#page-7-0) [6.3,](#page-11-0) and [7.1.](#page-11-1) In this 83 sense, kernels are important links between deterministic <sup>84</sup> and probabilistic models.

A strictly positive definite kernel K is called *stationary* if  $K(x, y) \equiv K(x - y)$ . According to Bochner's theorem  $[27]$ , K is the Fourier transform of a positive and bounded measure  $F$ , that is

$$
K(\boldsymbol{x}-\boldsymbol{y})=\int_{\mathbb{R}^d} \mathrm{e}^{\mathrm{i}(\boldsymbol{x}-\boldsymbol{y},\boldsymbol{\omega})} F(\mathrm{d} \boldsymbol{\omega}), \qquad \boldsymbol{x},\boldsymbol{y} \in \mathbb{R}^d.
$$

Here,  $(\cdot, \cdot)$  is the inner product in  $\mathbb{R}^d$  and i is the unit complex number. Fourier inversion is possible when  $K$ is absolutely integrable, in which case we call the Fourier transform  $\widehat{K}$  its *spectral density*. We note that  $\widehat{K}$  is nonnegative and integrable. Furthermore, most of the paper assumes stationarity and isotropy for the covariance function,  $K$ , so that

<span id="page-2-1"></span>(5) 
$$
Cov(Z(\boldsymbol{x}), Z(\boldsymbol{y})) = K(\boldsymbol{x} - \boldsymbol{y}) = \sigma^2 \varphi(||\boldsymbol{x} - \boldsymbol{y}||),
$$

for  $x, y \in \mathbb{R}^d$  and  $\|\cdot\|$  denoting the Euclidean distance. Here, we assume  $\varphi$  to be continuous with  $\varphi(0) = 1$ . Throughout, we shall equivalently call  $\varphi$  a *function* or a *correlation function*, the last as a shortcut to  $\varphi(\|\cdot\|)$ . Hence, the parameter  $\sigma^2 > 0$  is the variance of  $Z(\mathbf{x})$ , for all  $x \in \mathbb{R}^d$ . Let  $\Phi_d$  denote the class of such functions  $\varphi$  inducing a covariance function K through the iden-tity [\(5\)](#page-2-1) *i.e.*  $\Phi_d$  is the class of continuous isotropic correlation functions defined on  $\mathbb{R}^d$ . Such functions have a precise integral representation according to Schoenberg <sub>103</sub>  $[143]$ , given by

(6) 
$$
\varphi(x) = \int_0^\infty \Omega_d(rx) F_d(\mathrm{d}r), \qquad x \ge 0,
$$

with  $F_d$  being a probability measure and

(7) 
$$
\Omega_d(x) = \Gamma(d/2) \left(\frac{2}{x}\right)^{d/2-1} J_{d/2-1}(x), \qquad x \ge 0,
$$

with  $\Gamma(\cdot)$  the gamma function and  $J_{\nu}$  the Bessel function of the first kind of order  $\nu > 0$  [\[119,](#page-20-2) formula 10.2.2]. For <sub>104</sub> a member  $\varphi$  of the class  $\Phi_d$ , we can use that its d-variate Fourier transform of  $\varphi(\|\boldsymbol{x}-\boldsymbol{y}\|)$  is isotropic again, and <sub>106</sub> therefore reducible to a scalar integral formula (8)

$$
\widehat{\varphi}(z) = \frac{z^{1-d/2}}{(2\pi)^{d/2}} \int_0^\infty u^{d/2} J_{d/2-1}(uz) \varphi(u) \mathrm{d}u, \, z \ge 0,
$$

<sup>85</sup> defining its d-variate *isotropic spectral density*, and we assume this integral to exist. If the denominator  $(2\pi)^{d/2}$ 86 <sup>87</sup> is omitted, the same formula holds for the inverse ra-88 dial Fourier transform. Throughout, we write  $\Phi_{\infty}$  for <sup>89</sup>  $\bigcap_{d\geq 1} \Phi_d$ , the class of functions  $\varphi$  inducing positive def- $90$  inite radial functions on every d-dimensional Euclidean 91 space. Hence,  $\varphi \in \Phi_d$  if and only if  $\varphi(\|\cdot\|)$  is a correla-92 tion function in  $\mathbb{R}^d$ .

#### **2. THE MATÉRN MODEL**

<span id="page-2-2"></span><span id="page-2-0"></span>The *Matérn model*,  $\mathcal{M}_{\nu,\alpha}$ , is defined as [\[149\]](#page-21-3)

$$
(9) \qquad \mathcal{M}_{\nu,\alpha}(x) = \frac{2^{1-\nu}}{\Gamma(\nu)} \left(\frac{x}{\alpha}\right)^{\nu} \mathcal{K}_{\nu}\left(\frac{x}{\alpha}\right), \qquad x \ge 0,
$$

93 with  $\alpha > 0$  the *scale* parameter,  $\nu > 0$  the *smoothness* pa-94 rameters, and  $\mathcal{K}_{\nu}$  a modified Bessel function of the sec-95 ond kind of order  $\nu$  [\[2,](#page-17-0) 9.6.22]. It can be verified that 96  $\mathcal{M}_{\nu,\alpha}(0) = 1$ , so that [\(9\)](#page-2-2) is a correlation function. Argu-97 ments in Stein [\[149,](#page-21-3) p48] show that  $\mathcal{M}_{\nu,\alpha}$  belongs to the 98 class  $\Phi_{\infty}$ . The function  $\sigma^2 \mathcal{M}_{\nu,\alpha}$  will be termed *Matérn* <sup>99</sup> *covariance function*, and  $\sigma^2 > 0$  will denote the variance <sup>100</sup> of the associated Gaussian random field.

The importance of the Matérn class stems from the parameter  $\nu$  that controls the differentiability of the sample paths of the associated Gaussian field. Specifically, for any positive integer  $k$ , the sample paths of a Gaussian field  $Z$  on  $\mathbb{R}^d$  with Matérn correlation function are  $k$ -times mean square differentiable (in any direction) if and only if  $\nu > k$ . Also, a rescaled version of the Matérn correlation function converges to the Gaussian or squared exponential kernel as  $\nu \rightarrow \infty$ , that is

(10) 
$$
\mathcal{M}_{\nu,\alpha/(2\sqrt{\nu})}(x) \xrightarrow[\nu \to \infty]{} \exp(-x^2/\alpha^2), \qquad x \ge 0,
$$

with convergence being uniform on any compact set of 102  $\mathbb{R}^d$ . For this reason, the parametrisation  $\mathcal{M}_{\nu,\alpha/(2\sqrt{\nu})}$  is sometimes also adopted [\[170\]](#page-21-4).

When  $\nu = k + 1/2$ , for k a nonnegative integer, the Matérn correlation function simplifies into the product of a negative exponential correlation function with a polynomial of order k. For instance,  $\mathcal{M}_{1/2,1}(x) = \exp(-x)$  and  $\mathcal{M}_{3/2,1}(x) = \exp(-x)(1+x)$ . In general, (11)

$$
\mathcal{M}_{k+1/2,1}(x) = \exp(-x) \sum_{i=0}^{k} \frac{(k+i)!}{2k!} {k \choose i} (2x)^{k-i}
$$

for  $k \in \mathbb{N}_0$ . This simple algebraic form for the Matérn <sup>105</sup> correlation functions has undoubtedly contributed to the widespread popularity of the Matérn model.

<sup>107</sup> Now we are in a position to explore in detail the many <sup>108</sup> faces of the Matérn model. Section [3](#page-2-3) discusses maximum <sup>109</sup> likelihood estimation, Kriging prediction, and the screen-110 ing effect, while Section [4](#page-5-0) explores an SPDE characterisa-tion of the Matérn model. Section [5](#page-8-0) discusses the Matérn model as a building block to more sophisticated models, while Section  $6$  views the scientific landscape through the lens of the Matérn model, with special emphasis on nu-<sup>115</sup> merical analysis, probability theory and machine learn-<sup>116</sup> ing. Section [7](#page-11-2) introduces some recently developed alternatives and generalisations of the Matérn model, while Section [8](#page-13-0) compares these alternative models in terms of <sup>119</sup> estimation, prediction, and the screening effect.

# <span id="page-2-3"></span>**3. ESTIMATION AND PREDICTION WITH THE MATÉRN MODEL**

Let  $D \subset \mathbb{R}^d$  be a subset of  $\mathbb{R}^d$ . Consider a set  $X_n =$  ${x_1, \ldots, x_n}$  of (distinct) locations in D, at which values  $\mathbf{Z}_n = (Z(\boldsymbol{x}_1), \dots, Z(\boldsymbol{x}_n))^\top$  of the Gaussian random field  $Z$ , defined in Section [1.1,](#page-1-1) are observed. An important problem concerns the *prediction* of values  $Z(x_0)$  at an unobserved location  $x_0 \in D \setminus X_n$ . Then an especially natural predictor for  $Z(x_0)$  is

<span id="page-2-4"></span>(12) 
$$
\widehat{Z}_n = \boldsymbol{c}_n^\top \boldsymbol{R}_n^{-1} \boldsymbol{Z}_n
$$

with the vector  $[c_n]_i = K(x_0, x_i)$  and the *kernel matrix* <sup>121</sup>  $[\mathbf{R}_n]_{i,j} = K(\mathbf{x}_i, \mathbf{x}_j)$ . The predictor [\(12\)](#page-2-4) can be motivated

122 from multiple directions. Classically,  $(12)$  is motivated 166 123 as the best linear unbiased predictor (BLUP) for  $Z(\mathbf{x}_0)$ , 167 <sup>124</sup> and is often referred to as the *simple Kriging* predictor of 125  $Z(x_0)$  [\[42\]](#page-18-1). From a modern perspective, where the role 169 126 of unbiased estimation is increasingly questioned, we can 170 127 motivate this choice using alternative optimality proper-171 <sup>128</sup> ties, including:

- 129 1. it is the expectation of  $Z(x_0)$  conditionally on the 130 realisation  $Z_n$ ;
- <sup>131</sup> 2. it is the optimal estimate (i.e. the Bayes act) for  $Z(x_0)$  based on the data-set  $\mathbf{Z}_n$ , under squared er-133 ror loss [\[117,](#page-20-3) Section 13.3];
- <sup>134</sup> 3. it yields the minimal RKHS norm interpolant of the <sup>135</sup> data evaluated at  $x_0$ , by Section [6.1;](#page-9-0)
- <sup>136</sup> 4. it is the algorithm for approximating  $Z(x_0)$  from  $Z_n$  that minimises the worst case error in the <sup>138</sup> sense of information-based complexity [\[117,](#page-20-3) Sec-<sup>139</sup> tion 10.2] and approximation theory (see Section  $140$  [6.1\)](#page-9-0),

<sup>141</sup> to name but a few. The Matérn model provides a natu- $_{142}$  ral setting to study the performance of [\(12\)](#page-2-4) if we sup- $143$  pose  $Z$  to have a stationary isotropic covariance function <sup>144</sup>  $\sigma^2 \mathcal{M}_{\nu,\alpha}$ . The crucial question of how to select suitable 145 values for the parameters  $\sigma$ ,  $\alpha$ ,  $\nu$  will be considered first,  $146$  in Section [3.1,](#page-3-0) and then the performance of  $(12)$  will be 147 studied in Section [3.2.](#page-4-0) The possibility of a direct exten-<sup>148</sup> sion of the Matérn model to more general domains, such <sup>149</sup> as manifolds and graphs, is discussed in Section [3.3.](#page-5-1)

#### <span id="page-3-0"></span><sup>150</sup> **3.1 Estimation Using Maximum Likelihood**

 Maximum likelihood (ML) and similar estimation methods are popular in this setting due to the availabil- ity of practical (inc. gradient-based) numerical methods for computation and the classical theory that underpins ML. On the other hand, implicit in the use of ML is that the statistical model is well-specified, and this judgement must be made on a case-by-case basis. To limit scope, we focus on ML estimation in the sequel. Our aim is to understand when the parameters of the Matérn model can be consistently estimated from data, and to understand the asymptotic distribution of the ML estimator. To this end, recall that the Gaussian log-likelihood function is

<span id="page-3-1"></span>(13) 
$$
\mathcal{L}_n(\boldsymbol{\theta}) = -\frac{1}{2} \left( \log(|\sigma^2 \boldsymbol{R}_n)| \right) + \frac{1}{\sigma^2} \boldsymbol{Z}_n^\top \boldsymbol{R}_n^{-1} \boldsymbol{Z}_n \right),
$$

up to an additive constant, with  $\boldsymbol{\theta} = (\nu, \alpha, \sigma^2)$ . The ML estimator is defined as

<span id="page-3-2"></span>(14) 
$$
\widehat{\boldsymbol{\theta}}_n = \operatorname*{argmax}_{\boldsymbol{\theta} \in \mathbb{R}^3_+} \mathcal{L}_n(\boldsymbol{\theta}).
$$

<sup>163</sup> The ML estimate for the variance parameter can be comto puted in closed-form as  $\hat{\sigma}_n^2 = \mathbf{Z}_n^{\top} \mathbf{R}_n^{-1} \mathbf{Z}_n/n$ ; plugging<br>the expression into (13) reduces the numerical problem to this expression into  $(13)$  reduces the numerical problem to

<sup>166</sup> optimisation of a so-called *concentrated likelihood* over  $\mathbb{R}^2_+$ . However, maximizing the log-(concentrated) likelihood requires a nonlinear optimisation problem to be solved, for which numerical methods must be used; see Section [4.3.](#page-7-1)

The performance of ML estimation has been studied <sup>172</sup> principally in two different asymptotic limits. Under *fixed domain asymptotics*, the sampling domain  $D$  is bounded and the set of sampled locations  $X_n$  becomes increasingly dense in D. Under *increasing domain asymptotics*, the domain  $D$  grows with the number  $n$  of observed data, and the distance between any two sampled locations is bounded away from zero. Zhang and Zimmerman [\[181\]](#page-22-0) note that the peformance of the ML estimator can be quite different under these two frameworks, as will now be discussed.

<span id="page-3-3"></span>*3.1.1 Increasing Domain Asymptotics.* Mardia and Marshall [\[108\]](#page-20-4) make use of increasing domain asymptotics to establish, under mild regularity conditions, that the ML estimator is *strongly consistent*, meaning that  $\hat{\theta}_n \stackrel{a.s.}{\longrightarrow} \theta_0$  for the *true* parameter  $\psi_0$ . Furthermore, they establish that the ML estimator is *asymptotically normal*, meaning that

(15) 
$$
\boldsymbol{F}^{1/2}(\boldsymbol{\theta}_0)(\widehat{\boldsymbol{\theta}}_n-\boldsymbol{\theta}_0)\stackrel{d}{\longrightarrow}\mathcal{N}(\boldsymbol{0},\boldsymbol{I})
$$

where  $\boldsymbol{F}(\boldsymbol{\theta}) = -E[\mathcal{L}_n''(\boldsymbol{\theta})]$  is the Fisher information matrix, whose entries are

$$
F(\boldsymbol{\theta})_{i,j} = \frac{1}{2} \text{tr} \left( \frac{\mathrm{d} \mathbf{\Sigma}_n}{\mathrm{d} \boldsymbol{\theta}_i} \mathbf{\Sigma}_n^{-1} \frac{\mathrm{d} \mathbf{\Sigma}_n}{\mathrm{d} \boldsymbol{\theta}_j} \mathbf{\Sigma}_n^{-1} \right),\,
$$

<sup>182</sup> and  $\Sigma_n = \sigma^2 R_n$ . Although our focus is on the Matérn 183 model, we note that these kind of asymptotic results hold <sup>184</sup> for any parametric correlation function obeying particu-<sup>185</sup> lar regularity conditions that are stated in terms of eigen-<sup>186</sup> value conditions on the correlation matrix and its derivatives  $[108]$ , thought these may not be easy to verify in <sup>188</sup> general (see for instance Shaby and Ruppert [\[145\]](#page-21-5), for the exponential case). Generally speaking, as long as the spa-<sup>190</sup> tial extent of the sampling region is large compared with <sup>191</sup> the range of dependence of the random field, increasing-<sup>192</sup> domain asymptotics provide a very accurate description 193 of the behavior of the ML estimate [\[181,](#page-22-0) [145,](#page-21-5) [83\]](#page-19-2).

*3.1.2 Fixed Domain Asymptotics.* Zhang [\[180\]](#page-22-1) considered ML estimation for the Matérn model under fixed domain asymptotics, proving that when the smoothness parameter  $\nu$  is known and fixed, none of the parameters  $\sigma^2$ and  $\alpha$  can be estimated consistently when  $d = 1, 2, 3$ . Instead, only the parameter

(16) 
$$
\text{micro}_{\mathcal{M}} = \sigma^2 / \alpha^{2\nu},
$$

sometimes called *microergodic* parameter [\[181,](#page-22-0) [149\]](#page-21-3), can be consistently estimated. This is a consequence of the equivalence of the two corresponding Gaussian measures,

that we denote with  $P(\sigma_i^2 \mathcal{M}_{\nu,\alpha_i})$ , with  $i = 0, 1$ . In particular, for any bounded infinite set  $D \subset \mathbb{R}^d$ ,  $d = 1, 2, 3$ ,  $P(\sigma_0^2 M_{\nu,\alpha_0})$  is equivalent to  $P(\sigma_1^2 M_{\nu,\alpha_1})$  on the paths of  $Z(\mathbf{x}), \mathbf{x} \in D$ , if and only if

<span id="page-4-1"></span>(17) 
$$
\sigma_0^2/\alpha_0^{2\nu} = \sigma_1^2/\alpha_1^{2\nu}.
$$

194 In contrast, for  $d \geq 5$ , Anderes [\[7\]](#page-18-2) proved the orthogonal-<sup>195</sup> ity of two Gaussian measures with different Matérn co-<sup>196</sup> variance functions and hence, in this case, all the param-<sup>197</sup> eters can be consistently estimated under fixed-domain 198 asymptotics. The case  $d = 4$  has been recently studied in 199 Bolin and Kirchner [\[30\]](#page-18-3).

Asymptotic results associated with ML estimation of the microergodic parameter, again for a fixed known smoothness parameter  $\nu$ , can be found in Zhang [\[180\]](#page-22-1), and later on in Kaufman and Shaby [\[83\]](#page-19-2). In particular, for a zero mean Gaussian field defined on a bounded infinite set  $D \subset \mathbb{R}^d$ ,  $d = 1, 2, 3$ , with a Matérn covariance function  $\sigma_0^2 M_{\nu,\alpha_0}$  the ML estimator  $\hat{\sigma}_n^2 / \hat{\alpha}_n^{2\nu}$  of the microergodic parameter is strongly consistent, i.e.,

$$
\hat{\sigma}_n^2/\hat{\alpha}_n^{2\nu}\xrightarrow{a.s.}\sigma_0^2/\alpha_0^{2\nu},
$$

and its asymptotic distribution is given by

$$
\sqrt{n}(\hat{\sigma}_n^2/\hat{\alpha}_n^{2\nu} - \sigma_0^2/\alpha_0^{2\nu}) \stackrel{d}{\longrightarrow} \mathcal{N}(0, 2(\sigma_0^2/\alpha_0^{2\nu})^2).
$$

 Generally speaking, when the range of dependence of the random field is large with respect to the spatial extent of the sampling region, fixed domain asymptotics provide a very accurate description of the behavior of the ML es- timate of the microergodic parameter [\[83\]](#page-19-2). Extensions of <sub>205</sub> these results to the case where  $Z$  is observed with Gaus- sian errors can be found in Tang et al. [\[157\]](#page-21-6), while re- sults for a space-time version of the Matérn model can be found in Ip and Li [\[76\]](#page-19-3) and Faouzi et al. [\[53\]](#page-19-4). Finally we highlight that the efficient estimation of the microergodic parameter assuming the smoothness parameter unknown is still an open problem; some promising results in this  $_{212}$  direction can be found in Loh et al. [\[106\]](#page-20-5).

213 A recent article  $[105]$  relaxes the conditions imposed  $_{233}$ 214 by [\[157\]](#page-21-6) where the latter assumes that  $\nu$  is known, in <sub>234</sub> 215 concert with some technical assumptions. Recent contri-<sup>216</sup> butions deal with Bayesian fixed domain asymptotics for 217 Matérn Gaussian random fields, and we mention  $[95]$  and <sub>237</sub> 218 more recently [\[96\]](#page-20-8).

## <span id="page-4-0"></span><sup>219</sup> **3.2 Prediction and the Screening Effect**

220 The equivalence of Gaussian measures within the  $_{241}$ 221 Matérn class has consequences for prediction of  $Z(\mathbf{x}_0)$  at <sub>242</sub> 222 an unobserved location  $x_0 \in D \setminus X_n$ ; these consequences <sub>243</sub> 223 will now be discussed. In what follows,  $\nu$  is supposed  $_{244}$ 224 known and fixed, and we consider the setting where  $\sigma_{245}$ 225 and  $\alpha$  are *misspecified*. That is, we suppose Z is a Gaus- <sub>246</sub> <sup>226</sup> sian field with Matérn covariance  $\sigma_0^2 M_{\nu,\alpha_0}$ , and we con-227 sider the performance of the predictor  $(12)$  when a Matérn

228 model  $\sigma_1^2 M_{\nu,\alpha_1}$  is used. This situation is typical, since the true parameters  $\sigma_0$  and  $\alpha_0$  of the data-generating process will be unknown in general. Our theoretical setting will be fixed domain asymptotics.

Note, first, that [\(12\)](#page-2-4) does not depend on the value of  $\sigma_1$ , but does depend on the value of the parameter  $\alpha_1$  (and the parameter  $\nu$ , but this parameter is fixed). This dependence will be emphasised using the notation  $c_n(\alpha_1)$  and  $\bm{R}_n(\alpha_1)$ . Under the Gaussian measure  $P(\sigma_0^2 \mathcal{M}_{\nu,\alpha_0})$  associated with the *true* model  $\sigma_0^2 M_{\nu,\alpha_0}$ , the mean squared error of the predictor  $\overline{Z}_n(\alpha_1)$  is given by

$$
\begin{aligned}\n&\text{VAR}_{\alpha_0, \sigma_0^2} \left[ \widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0) \right] \\
&= \sigma_0^2 \Big( 1 - 2 \boldsymbol{c}_n(\alpha_1)^\top \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{c}_n(\alpha_0) \\
&\quad + \boldsymbol{c}_n(\alpha_1)^\top \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{R}_n(\alpha_0) \boldsymbol{R}_n(\alpha_1)^{-1} \boldsymbol{c}_n(\alpha_1) \Big),\n\end{aligned}
$$

while if there is no misspecification then the previous expression reduces to

(18) 
$$
\operatorname{VAR}_{\alpha_0, \sigma_0^2} \left[ \widehat{Z}_n(\alpha_0) - Z(\boldsymbol{x}_0) \right] = \sigma_0^2 \big( 1 - \boldsymbol{c}_n(\alpha_0)^\top \boldsymbol{R}_n^{-1}(\alpha_0) \boldsymbol{c}_n(\alpha_0) \big).
$$

Under regularity conditions, and for fixed domain asymptotics, Stein [\[147\]](#page-21-7) shows that both asymptotically efficient prediction and asymptotically correct estimation of prediction variance hold when the two Gaussian measures  $P(\sigma_i^2 \mathcal{M}_{\nu,\alpha_i})$ ,  $i = 0,1$  are equivalent, *i.e.* [\(17\)](#page-4-1). Specifically,

<span id="page-4-2"></span>(19) 
$$
\frac{\text{VAR}_{\sigma_0^2, \alpha_0} \left[\widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0)\right]}{\text{VAR}_{\sigma_0^2, \alpha_0} \left[\widehat{Z}_n(\alpha_0) - Z(\boldsymbol{x}_0)\right]} \xrightarrow{a.s.} 1
$$

and

<span id="page-4-3"></span>(20) 
$$
\frac{\text{VAR}_{\sigma_1^2,\alpha_1}[\widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0)]}{\text{VAR}_{\sigma_0^2,\alpha_0}[\widehat{Z}_n(\alpha_1) - Z(\boldsymbol{x}_0)]} \xrightarrow{a.s.} 1.
$$

 $232$  The implication of  $(19)$  is that, under the true model, if the correct value of  $\nu$  is used, any value of  $\alpha_1$  will give asymptotic efficiency. The implication of  $(20)$  is stronger and guarantees that using the misspecified predictor un-<sup>236</sup> der the correct and misspecified models is asymptotically equivalent from mean squared error point of view. Note <sup>238</sup> that these kind of results does not consider the uncertainty <sup>239</sup> associated with the covariance parameters of the misspec- $_{240}$  ified model. Kaufman and Shaby [\[83\]](#page-19-2) show that [\(20\)](#page-4-3) still holds by considering the ML estimator of the variance <sup>242</sup>  $\hat{\sigma}_n^2 = \mathbf{Z}_n^{\top} \mathbf{R}_n^{-1}(\alpha_1) \mathbf{Z}_n/n$  in place  $\sigma_1^2$ .

Conditions of equivalence of two Gaussian measures based on a space-time  $[76]$  and bivariate  $[13]$  version of the Matérn model have also been established. Next, we consider a practically important aspect of prediction; the co-called screening effect.

 *Screening Effect.* The *screening effect* refers to the phe- nomenon where the predictor  $(12)$  depends almost ex- clusively on those observations that are located nearest to the predictand [\[150\]](#page-21-8). As such, the screening effect is an important tool that can be used to mitigate the com- putational burden of evaluating  $(12)$  in the presence of big datasets. This issue has traditionally been an impor- tant subject in geostatistics [\[110,](#page-20-9) [111,](#page-20-10) [112,](#page-20-11) [39\]](#page-18-5). Indeed, Matheron [\[110,](#page-20-9) [111\]](#page-20-10), in the School of Geostatistics at the Ecole des Mines, developed a first formalisation of screening effect, referring to situations where the observa- tions located far from the predictand receive a zero krig- ing weight. Matheron's definition has a direct connection with the Markov property on the real line, which happens when kriging is performed under the exponential model 263 (indeed,  $\mathcal{M}_{1/2,\alpha}$ ).

M. Stein [\[149,](#page-21-3) [150,](#page-21-8) [152,](#page-21-9) [153\]](#page-21-10) adopts an alternative <sub>274</sub> definition of the screening effect that will now be de- $_{275}$ scribed. Let Z be a mean-square continuous, zero mean  $_{276}$ and weakly stationary Gaussian random field on  $\mathbb{R}^d$ . Let  $e(X_n)$  be the error of the predictor [\(12\)](#page-2-4) of  $Z(\mathbf{x}_0)$  based on  $_{278}$  $\mathbf{Z}_n$ . Two choices for the set  $X_n$  of observation locations <sub>279</sub> will be considered, and to this end we let  $F_{\epsilon}$ ,  $N_{\epsilon}$  be sets, <sub>280</sub> indexed by  $\epsilon > 0$ , such that  $N_{\epsilon}$  contains the nearest obser-  $_{281}$ vations to the predictand, and  $F_{\epsilon}$  the furthest observations. 282 Then Stein  $[150]$  says that  $N_e$  *asymptotically screens out* 283  $F_{\epsilon}$  when

<span id="page-5-2"></span>(21) 
$$
\lim_{\epsilon \downarrow 0} \frac{\mathbb{E} e(N_{\epsilon} \cup F_{\epsilon})^2}{\mathbb{E} e(N_{\epsilon})^2} = 1.
$$

 A thorough discussion of the implications of this defini- tion can be found in Porcu et al.  $[130]$ , where nontrivial  $289$  differences between fixed domain and increasing domain asymptotics are reported.

The spatial configuration of the sampling point  $X_n$  determines whether the screening effect will hold. Porcu et al. [\[130\]](#page-21-11) refer to a *regular scheme* as one for which  $F_{\epsilon} = {\epsilon(\mathbf{x}_0 + j)}$ , for  $j \in \mathbb{Z}^d$  and  $N_{\epsilon}$  being the restriction of  $F_{\epsilon}$  to some fixed region with  $x_0$  in its interior, assuming  $x_0 \notin \mathbb{Z}^d$ . For regular schemes, Stein [\[150\]](#page-21-8) established  $(21)$  whenever *the spectrum*  $\hat{K}$  *varies regularly at infinity [\[26\]](#page-18-6) in every direction with a common index of variation* [quoted from [130\]](#page-21-11). However, this condition may not be useful for space-time processes, where differentiability properties in the space and time coordinates are not <sub>303</sub> necessarily identical. To overcome such a problem, we instead consider an *irregular scheme*: for  $x_1, \ldots, x_n$  being <sub>305</sub> distinct nonzero elements of  $\mathbb{R}^d$ ,  $\pmb{y}_1, \ldots, \pmb{y}_N$  distinct elements of  $\mathbb{R}^d$ ,  $x_0 = 0 \in \mathbb{R}^d$  and  $y_0 \in \mathbb{R}^d$  being nonzero, we have  $N_{\epsilon} = \{\epsilon x_1, \ldots, \epsilon x_n\}$  and  $F_{\epsilon} = \{y_0 + \epsilon y_1, \ldots, y_0 + \epsilon y_n\}$  $\{\epsilon y_N\}$ . The *Stein hypothesis* [termed in [130\]](#page-21-11)

<span id="page-5-3"></span>(22) 
$$
\forall R > 0
$$
,  $\lim_{\|\boldsymbol{\omega}\| \to \infty} \sup_{\|\boldsymbol{\tau}\| < R} \left| \frac{\hat{K}(\boldsymbol{\omega} + \boldsymbol{\tau})}{\hat{K}(\boldsymbol{\omega})} - 1 \right| = 0$ ,

provides a sufficient condition for the screening effect in this setting (under some mild additional conditions on  $K$ and  $N<sub>\epsilon</sub>$ ), which can be verified in dimensions  $d = 1$  and  $d = 2$  for mean-square continuous but non-differentiable random fields, for some specific designs  $N_e$  [\[152\]](#page-21-9). The Matérn model with  $K = M_{\alpha,\nu}$  admits a simple expression for its spectrum [\[2,](#page-17-0) 11.4.44]:

<span id="page-5-4"></span>(23) 
$$
\widehat{\mathcal{M}}_{\nu,\alpha}(z) = \frac{\Gamma(\nu + d/2)}{\pi^{d/2} \Gamma(\nu)} \frac{\alpha^d}{(1 + \alpha^2 z^2)^{\nu + d/2}}, \ z \ge 0,
$$

from which  $(22)$  can be verified.

<sup>269</sup> The screening effect can thus be established for the Matérn model, under both regular and irregular schemes, justifying the use of "local" approximations to the predic-272 tor  $(12)$ .

#### <span id="page-5-1"></span><sup>273</sup> **3.3 Matérn on Manifolds and Graphs**

Let  $M$  be a general manifold. A pragmatic question is whether the Matérn correlation function  $(9)$  can be composed with a suitable metric  $q$ , defined on the manifold, to preserve positive definiteness over  $M$ . For the case of the sphere, a natural metric is the geodesic distance; the length of the arc connecting any pair of points located over the spherical shell. For this metric,  $(x, y) \mapsto$  $\mathcal{M}_{\nu,\alpha}(g(x,y))$  is a correlation function only for  $0 < \nu \leq$  $1/2$  [\[61\]](#page-19-5). This limitation is emphasised in Alegría et al. [\[3\]](#page-17-1), who propose the  $\mathcal F$  family, a model that is valid on <sup>284</sup> the sphere, and having the same properties as the Matérn <sup>285</sup> function in terms of mean-square differentiability and <sup>286</sup> fractal dimension. The Matérn function on other general  $287$  manifolds has been studied by Li et al. [\[98\]](#page-20-12). Guinness and Fuentes [\[68\]](#page-19-6) propose a spectral expansion to define a covariance function that mimics the Matérn model, but this 290 construction is criticised by Lindgren et al.  $[101]$  as being incorrect as the spectral expansion does not reproduce the <sup>292</sup> same properties of the Matérn model.

Unfortunately, it seems that the limited applicability of the Matérn model on any space that is not a flat surface extends to more abstract settings as well. An elegant iso-<sup>296</sup> metric embedding argument in Anderes et al. [\[8\]](#page-18-7) proves that the restriction  $0 < \nu \leq 1/2$  is required when the input space is a graph with Euclidean edges. A more general argument in Menegatto et al.  $[113]$  proves that the same <sup>300</sup> restriction is inherited for a general quasi metric space en-<sup>301</sup> dowed with a geodesic metric. The notable effort by Bolin 302 and Kirchner [\[29\]](#page-18-8) provides a model that is once differentiable over metric graphs. It is reasonable to conclude that some form of the SPDE approach, which we discuss next in Section  $4.2$ , is needed in general to extend the Matérn <sup>306</sup> model to a general manifold.

# <span id="page-5-0"></span>**4. THE MATÉRN MODEL IN COMPUTATIONAL STATISTICS**

<sup>307</sup> This section explores the interaction of the Matérn <sup>308</sup> model with computational statistics, starting with numer-<sup>309</sup> ical methods for *implementation* of the Matérn model  $310$  (Sections [4.1,](#page-6-1) [4.2](#page-6-0) and [4.3\)](#page-7-1), and then turning to uses of  $332$ <sup>311</sup> the Matérn model to *facilitate* numerical computation it- $312$  self (Section [4.4\)](#page-7-0).

#### <span id="page-6-1"></span><sup>313</sup> **4.1 Implementation as an SDE**

The Matérn model admits a *state space* representation as an SDE, which enables efficient computational techniques from the signal processing literature to be employed for simulation, estimation and prediction. Indeed, focusing on dimension  $d = 1$ , and letting

$$
\mathbf{Z}(x) = (Z, dZ/dx, \dots, d^k Z/dx^k),
$$

the Matérn model  $\mathcal{M}_{\nu,\alpha}$  with  $\nu = k + 1/2$  admits the 344 characterisation

$$
d\mathbf{Z} = \begin{pmatrix} 0 & 1 \\ \vdots & \vdots \\ 0 & 1 \\ -a_0 - a_1 \dots - a_{k-1} \end{pmatrix} \mathbf{Z} dx + \begin{pmatrix} 0 \\ \vdots \\ 0 \\ 1 \end{pmatrix} d\mathcal{W}
$$

<sup>314</sup> where  $a_i = {}_{k+1}C_i \cdot \alpha^{-k-1+i}$ , the  $C$  are binomial coef- ficients, and  $W(x)$  represents a zero-mean white noise 316 process on  $x \in \mathbb{R}$  [\[72\]](#page-19-7). The advantage of state space for-317 mulations is that both estimation and prediction can be performed in a *single pass* through the data, at linear O(n) cost, using familiar Kalman updating equations as 320 described in Sarkka et al. [\[136\]](#page-21-12) and in further detail in Chapter 6 of Hennig et al. [\[74\]](#page-19-8). Similar characterisations for higher dimensions, including spatio-temporal versions 323 of the Matérn model, can be found in Sarkka et al. [\[136\]](#page-21-12), though we note these retain linear complexity only in the number of time steps; complexity is cubic in the size of the spatial grid. The SPDE approach can offer a solution in this respect, and we discuss this next.

# <span id="page-6-0"></span><sup>328</sup> **4.2 Implementation as an SPDE**

A major reason for the continued popularity of the Matérn model is the availability of efficient and scalable numerical methods for simulation, due in large part to Lindgren et al. [\[102\]](#page-20-15). These authors consider the SPDE

<span id="page-6-2"></span>
$$
(24) \qquad (\alpha^{-2}-\Delta)^{\gamma/2}Z(\boldsymbol{x})=\mathcal{W}(\boldsymbol{x}), \qquad \boldsymbol{x}\in\mathbb{R}^d,
$$

where  $\alpha > 0$ ,  $\Delta$  is the Laplacian, and W is a Gaussian white noise on  $\mathbb{R}^d$ , so that Cov  $(\mathcal{W}(A_1), \mathcal{W}(A_2)) = |A_1 \cap$  $A_2$ , where  $A_i$  are subsets of  $\mathbb{R}^d$ ,  $i = 1, 2$ , and where  $|\cdot|$ is the volume integral. Whittle [\[168\]](#page-21-13) and Whittle [\[169\]](#page-21-14) proved that the solution to  $(24)$  is a Gaussian field with Matérn covariance  $\sigma^2 \mathcal{M}_{\nu,\alpha}$  with parameters  $\alpha$  (as before) and

$$
\sigma^2 = \frac{\Gamma(\nu)\alpha^{2\nu}}{\Gamma(\nu + d/2)(4\pi)^{d/2}}, \qquad \nu = \gamma - d/2.
$$

329 This perspective offers two insights; first, tools developed 375 330 for the numerical approximation of SPDEs can be brought 376 331 to bear on the Matérn model, and second, there is a clear 377

path to generalise the definition of the Matérn model to any (planar or non planar) manifold on which the analo-<sup>334</sup> gous SPDE may be defined. (For example, Jansson et al. 335 [\[77\]](#page-19-9) take this perspective to generalise the Matérn model 336 to the sphere  $\tilde{\mathbb{S}}^d$ .)

To provide a computationally convenient approximation to  $(24)$ , Lindgren et al. [\[102\]](#page-20-15) considered the weak solution to  $(24)$  and approximation of the weak solution using basis functions with compact support over a com- $_{341}$  pact domain  $\Omega \subset \mathbb{R}^d$  (specifically, a *Galerkin* approxima-<sup>342</sup> tion using finite element basis functions was used). As <sup>343</sup> a result, the authors establish a formal route to approximation of the random field Z with a *Gauss–Markov* ran-<sup>345</sup> dom field having a *sparse* precision matrix. Sparse matrix <sup>346</sup> algebra enables fast simulation of realisations from the <sup>347</sup> Matérn random field, and fast evaluation of the likelihood  $348$  [\(13\)](#page-3-1) (albeit not fast evaluation of the gradient of the like-<sup>349</sup> lihood).

350 The choice of domain  $\Omega$  introduces boundary effects  $351$  which must be carefully mitigated. Khristenko et al.  $[86]$ , Brown et al. [\[35\]](#page-18-9) provide a solution for the case where  $\gamma$  is an integer; the non-integer case is considered in Bolin and Kirchner [\[29\]](#page-18-8). The extension of the Matérn field based on 355 SPDEs to space-time is provided by Cameletti et al. [\[37\]](#page-18-10) and subsequently by Bakka et al.  $[14]$ , Clarotto et al.  $[40]$ , while the multivariate Matérn case has been explored in Bolin and Wallin [\[32\]](#page-18-13). Alternative approximations based <sup>359</sup> on Galerkin methods on manifolds have been provided 360 by Lang and Pereira [\[90\]](#page-20-16). An interesting approach that <sup>361</sup> allows working on manifolds with huge datasets is pro- $_{362}$  posed by Pereira et al. [\[122\]](#page-20-17). The interest in this literature 363 is dual. On the one hand, the technical aspects related to <sup>364</sup> the finite dimensional representation of Gaussian random <sup>365</sup> fields are extremely interesting *per se*. On the other hand, <sup>366</sup> this group of authors is actually driven by providing tools for efficient computation. This is witnessed by the relevant existing R packages (R-INLA, inlabru, and rSPDE <sup>369</sup> for instance) and we refer to the review of Lindgren et al. <sup>370</sup> [\[101\]](#page-20-13).

Sanz-Alonso and Yang [\[134\]](#page-21-15) attempt to explain the trade-off between accuracy and scalability in numerical approximation of the Matérn model. Recall that, in the SPDE approach  $[102]$ , Z in  $(24)$  is numerically approximated using a Gaussian process

<span id="page-6-3"></span>(25) 
$$
Z_{\delta}(\boldsymbol{x}) = \sum_{k=1}^{n_{\delta}} \omega_k \epsilon_k(\boldsymbol{x}), \qquad \boldsymbol{x} \in \Omega,
$$

 $371$  where  $\epsilon_k$  are finite element basis functions and the vec-<sup>372</sup> tor  $\boldsymbol{\omega} = (\omega_1, \dots, \omega_{n_{\delta}})^\top$  is multivariate Gaussian with zero 373 mean and with a sparse precision matrix. The accuracy 374 of the approximation  $Z_{\delta}$  is dependent on (a) the compact support of the finite elements basis functions, (b) boundary effects due to the domain  $\Omega$ , and (c) by the mesh width  $\delta$  that determines the cardinality  $n_{\delta}$  in [\(25\)](#page-6-3). Most of the 378 earlier literature has considered [\(25\)](#page-6-3) with  $n_{\delta}$  proportional 431  $379$  to the sample size n of the dataset being modelled. Sanz-  $432$ 380 Alonso and Yang [\[134\]](#page-21-15) adopt a fixed domain asymptotic 433 381 approach to explain when  $n_{\delta} \ll n$  might be a legitimate 434 <sup>382</sup> strategy. To do so, they consider Gaussian process regres-<sup>383</sup> sion and work under the framework of Bayesian contrac-<sup>384</sup> tion rates. Their results provide justification for specific 385 scalings of  $n_{\delta}$  with  $n_{\delta} = o(n)$ , provided that the smooth-438 386 ness  $\nu$  is sufficiently high.

 A different path to SPDE and Gauss–Markov random  $388$  fields was recently taken in Sanz-Alonso and Yang [\[135\]](#page-21-16),  $441$  who adopt graph-based discretisations of SPDEs. This ap- proach can be well-suited to working with discrete and 391 unstructured point clouds, such as in machine learning 444 tasks where the data belong to an implicitly defined low- dimensional manifold. A second advantage of this ap- proach is that an explicit triangulation of the domain is not required.

#### <span id="page-7-1"></span><sup>396</sup> **4.3 Approximate Likelihood and the Matérn Model**

 In estimating the parameters of the Matérn model using ML [\(14\)](#page-3-2), numerical optimisation is required. Although generic optimisation routines can be used, an often better approach is to first construct a cheap approximation to the likelihood, which can then be more readily maximised. Indeed, approximate likelihoods are essential when deal- ing with large datasets, since the evaluation of  $(13)$  re- quires computing the inverse and the determinant of the correlation matrix, usually via the Cholesky decomposi-406 tion at complexity  $O(n^3)$  and storage cost  $O(n^2)$ .

 Perhaps the most successful approximation is *Vecchia's method* [\[164\]](#page-21-17), which has attracted a remarkable amount of attention in recent times [inc.  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$  $148, 46, 47, 66, 45$ ]. The Vecchia approximation can be used with any correlation 411 model and its basic idea is is to replace [\(13\)](#page-3-1) with a prod- uct of Gaussian conditional distributions, in which each conditional distribution involves only a small subset of the data. This approximation requires that the data are *or- dered* and the number m of 'previous' data on which to 416 condition is to be specified. Generally, larger  $m$  entails more accurate and computationally expensive approxima- tion, while the choice of ordering affects the accuracy of 419 the approximation [\[66\]](#page-19-12). The Vecchia method provides a sparse approximation to the Cholesky factor of the pre- cision matrix, such that the approximate likelihood can  $468$ <sup>422</sup> be computed in  $O(nm^3)$  time and with  $O(nm^2)$  storage cost. See the recent review of Katzfuss and Guinness [\[82\]](#page-19-13) for further detail. The Vecchia likelihood can be viewed as a specific instance of a more general class of estimation 426 methods called quasi- or composite likelihood  $[103, 163]$  $[103, 163]$  $[103, 163]$ <sub>473</sub> that have been widely used for the estimation of Gaussian 428 fields with the Matérn model  $[50, 24, 12]$  $[50, 24, 12]$  $[50, 24, 12]$  $[50, 24, 12]$  $[50, 24, 12]$ .

<sup>429</sup> An alternative method of mitigating the computational <sup>430</sup> burden of ML estimation is *covariance tapering* [\[57\]](#page-19-15). The

basic idea is to multiply the Matérn model with a compactly supported correlation function, resulting in a 'modified' Matérn model with compact support. This induces sparseness in the associated covariance matrix, so that algorithms for sparse matrices can be exploited for a computationally efficient evaluation of the Cholesky decomposition  $[57]$ . However, some authors  $[23, 21]$  $[23, 21]$  $[23, 21]$  suggest that tapering might be an obsolete approach in view of <sup>439</sup> the fact that flexible compactly supported models that include the Matérn model as a special case have been re-cently proposed; see Section [8.](#page-13-0) A comprehensive review of the likelihood approximations is beyond the scopes of this paper, so we refer the reader to Sun et al.  $[155]$  and Heaton et al.  $[73]$  for further detail.

# <span id="page-7-0"></span><sup>445</sup> **4.4 The Matérn Model** *for* **Bayesian Computation**

In the last decade there has been increasing interest in the use of kernel methods for solving PDEs. Consider a system

$$
\mathcal{A}u = f \qquad \text{in } \Omega
$$

$$
\mathcal{B}u = g \qquad \text{on } \partial\Omega
$$

specified by a differential equation involving  $A$  and f, at and initial or boundary conditions specified by  $\beta$  and  $q$ . 448 Dating back at least to Fasshauer [\[55\]](#page-19-17) in the determinis-<sup>449</sup> tic setting, and reinterpreted through a Bayesian lens by <sup>450</sup> authors such as Cockayne et al. [\[41\]](#page-18-20), one can seek an ap-451 proximation to the strong solution  $u : \Omega \to \mathbb{R}$  by mod-<sup>452</sup> elling u as *a priori* a Gaussian random field and condi-<sup>453</sup> tioning that field to satisfy the differential equation at lo-454 cations  $\{x_1, \ldots, x_m\} \subset \Omega$  and satisfy the boundary conditions at locations  $\{x_{m+1}, \ldots, x_n\} \subset \partial\Omega$ . The condi-<sup>456</sup> tional mean of this process coincides with the *symmetric* <sup>457</sup> *collocation* method introduced by Fasshauer [\[55\]](#page-19-17), which  $458$  we return to in Section [6.1,](#page-9-0) while the conditional variance provides probabilistic uncertainty quantification for the solution, expressing the uncertainty that remains as a <sup>461</sup> result of using only a finite computational budget. To implement these methods, one requires a Gaussian process <sup>463</sup> whose sample paths possess sufficient regularity for the operation of conditioning on the derivative  $Au$  to be welldefined. On the other hand, assuming excessive smooth-<sup>466</sup> ness could lead to over-confident uncertainty quantifi-<sup>467</sup> cation. One therefore requires a kernel with customisable smoothness, which can be adapted to the differential <sup>469</sup> equation at hand. The Matérn class satisfies this require-<sup>470</sup> ment, but is not alone in doing so; we continue discussion of this point in Section [7.](#page-11-2)

<sup>472</sup> A specific PDE that has received considerable recent attention in the Bayesian statistical community is the *Stein equation*, for which  $Au = c + p^{-1} \nabla \cdot (p \nabla u)$ , where p <sup>475</sup> is the probability density function of a posterior distribution of interest,  $f$  is a function whose posterior expectation we seek to compute, and  $c$  is a constant. If the Stein

 equation has a solution, then c *must* be the value of the posterior expectation we seek. This has motivated sev- eral efforts to numerically solve the Stein equation, as a 481 more direct alternative to first approximating  $p$  (for ex-  $532$  ample using Markov chain Monte Carlo) and then using 483 the approximation of p to approximate the expectation of interest. In this context kernel methods are typically used [\[118,](#page-20-19) [146\]](#page-21-21) and in particular the kernel should have smoothness that is two orders higher than that of the func- tion f whose expectation is of interest, since the Stein equation is a second-order PDE. The generalisation of the Stein equation to Riemannian manifolds was considered in [\[18\]](#page-18-21), who advocated for the use of kernels with cus- tomisable smoothness that reproduce Sobolev spaces of functions on the manifold, such as the (manifold gener- alisation of the) Matérn model. The connection between the Matérn model and Sobolev spaces is set out in Section 495  $6.1.$ 

# <span id="page-8-0"></span>**5. FLEXIBLE MODELLING WITH MATÉRN**

496 One might object that the Matérn model is insufficiently 550 flexible for many statistical applications, being limited to scalar-valued random fields that are stationary, isotropic and Gaussian. However, the Matérn model is also an im- portant building block for many more sophisticated mod- els, some of which will now be described. This is a rich literature, and our discussion is necessarily succinct; an  $_{503}$  extended version of this section can be found in [\[125\]](#page-20-20).

## **5.1 Scalar Valued Random Fields**

 Let us start by discussing models for scalar-valued ran- dom fields that build on the Matérn model. Note that one can trivially introduce non-zero mean functions into the Matérn model, or combine (additively or multiplicatively) kernels to obtain a potentially more expressive kernel; we will not dwell on either point.

 To relax the isotropy assumption of the Matérn model, [\[6\]](#page-18-22) consider scale mixtures that take into account pref-513 erential directions in which spatial dependence develops. On the other hand, the case of space-time models re- quires special treatment, and non-separable versions of the Matérn kernel are described in Gneiting [\[60\]](#page-19-18), Zas-tavnyi and Porcu [\[179\]](#page-22-2).

 The stationarity assumption was relaxed in a paramet- ric manner in Paciorek and Schervish [\[120\]](#page-20-21), and then in a nonparametric manner in Roininen et al. [\[133\]](#page-21-22). An attempt to strike a balance between the computational tractability of parametric models and the flexibility of nonparametric models was reported in Wilson et al. [\[171\]](#page-21-23), who proposed *input warping* to transform the inputs to the Matérn model using a neural network.

 The Gaussian assumption can be relaxed through *out-*<sup>527</sup> put warping, meaning transformation of the form  $\ddot{Z}(x) =$  $w(Z(x))$  where  $w(\cdot)$  is a nonlinear map from  $\mathbb{R}^d$  to  $\mathbb{R}^d$ .

The covariance function of  $\tilde{Z}$  will not be Matérn in general, when the covariance function of  $Z$  is Matérn, but if w is sufficiently regular then the smoothness properties of  $Z$  transfer to  $Z$ . The question of whether there exist non-Gaussian processes whose covariance function is nevertheless of Matérn class was answered positively in Åberg and Podgórski [\[1\]](#page-17-2). Yan and Genton [\[175\]](#page-22-3) have pro- posed *trans-Gaussian* random fields with Matérn covari- ance function. Bolin [\[28\]](#page-18-23) and subsequently Wallin and Bolin [\[165\]](#page-21-24) provided SPDE-based constructions for non- Gaussian Matérn fields. General classes of non-Gaussian fields with covariance  $g(\mathcal{M}_{\nu,\alpha})$ , for  $g(\cdot)$  a suitable func- tion that preserves the positive definiteness and smooth- ness properties of the Matérn model, have been provided for instance by Palacios and Steel  $[121]$ , Xua and Gen- ton [\[174\]](#page-22-4), Bevilacqua et al. [\[22\]](#page-18-24), Morales-Navarrete et al.  $[114]$ .

 An important extension of the Matérn model, which has received recent attention, is to random fields on spaces for which classical notions of smoothness are not well-defined. For example, Anderes et al. [\[8\]](#page-18-7) consider graphs with Euclidean edges, equipped with either the geodesic distance over the graph, or the resistance metric. Menegatto et al. [\[113\]](#page-20-14) provide a generalisation of this setting by considering quasi-metric spaces. Bolin et al.  $[31]$ adopt a different approach to build random fields with their covariance structure on metric graphs. Space-time version of the Matérn model, for graphs with Euclidean edges, have been considered by Tang and Zimmerman [\[156\]](#page-21-25) and Porcu et al. [\[129\]](#page-21-26). These efforts considerably extend the applicability of the Matérn model.

The Matérn covariance function decays exponentially with distance, which can be inappropriate for modelling processes that involve long memory. Several approaches have been developed to modify the tails of the Matérn cor- relation function while preserving many of its desirable characteristics; we describe these in Section [7.](#page-11-2)

[\[67\]](#page-19-19) considers Gaussian random fields defined for lat- $\bar{z}$  tices  $\bar{Z}^d$  with a covariance function that is the restriction 568 of the Matérn covariance to  $\mathbb{Z}^d$ . The resulting spectrum is smoothed version of the spectral density associated with the Matérn covariance. For this specific situation, the SPDE approximation can overestimate the scale,  $\alpha$ . Yet, it is not clear how this message extends to Gaussian fields that are continuously indexed in  $\mathbb{R}^d$ .

## **5.2 Vector-Valued Random Fields**

There has been a plethora of approaches related to multivariate spatial modeling, and the reader is referred to Genton and Kleiber [\[58\]](#page-19-1). Here, the isotropic covariance function  $\mathbf{K} : [0, \infty) \to \mathbb{R}^{p \times p}$  is matrix-valued. The elements on the diagonal,  $K_{ii}$ , are called *auto-covariance* functions, and the elements  $K_{ij}$ ,  $i \neq j$ , are called *crosscovariance* functions. Gneiting et al. [\[62\]](#page-19-20) proposed a multivariate Matérn model

<span id="page-8-1"></span>
$$
(26) \t K_{ij}(x) = \sigma_{ii}\sigma_{jj}\rho_{ij}\mathcal{M}_{\nu_{ij},\alpha_{ij}}(x), \t x \ge 0,
$$

<sup>575</sup> where  $\sigma_{ii}^2$  is the variance of  $Z_i$ , the *i*th component of a multivariate random field in  $\mathbb{R}^p$ , and  $\rho_{ij}$  is the collocated 577 correlation coefficient. There are restrictions on the pa- 630  $\epsilon_{578}$  rameters  $\nu_{ij}, \alpha_{ij}$  and  $\rho_{ij}$  required to ensure positive defi- niteness, and often the restrictions on the collocated cor- relations coefficients  $\rho_{ij}$  are rather strict. This last remark has motivated alternative approaches, and the reader is re- ferred to Apanasovich et al. [\[10\]](#page-18-26) and more recently to Emery et al. [\[52\]](#page-19-21). Extensions to multivariate space-time  $_{534}$ 584 Matérn structures have been provided by Allard et al. [\[5\]](#page-17-3) and through a technical approach by Porcu et al. [\[127\]](#page-20-24). Multivariate nonstationary Matérn functions have been proposed by Kleiber and Nychka [\[87\]](#page-20-25). Multivariate Matérn models with *dimple* effect have been studied by Alegría et al. [\[4\]](#page-17-4); a 'dimple' in a space-time covariance 590 model refers to the case when  $Cov(Z(\boldsymbol{x},t), Z(\boldsymbol{x}',t'))$  is  $\mathcal{L}_{\text{591}}$  bigger than  $\text{Cov}((Z(\bm{x}, t), Z(\bm{x}', t)),$  which requires spe-cial mathematical treatment.

 Multivariate Matérn modeling on graphs has been re- cently investigated in Dey et al. [\[49\]](#page-19-22), who propose a class of multivariate graphical Gaussian processes through *stitching*, a construction that gets multivariate covari- ance functions from the graph, and ensures process-level conditional independence between variables. When cou- pled with the Matérn model, stitching yields a multi- variate Gaussian process whose univariate components are Matérn Gaussian processes, and which agrees with process-level conditional independence as specified by the graphical model. Stitching can offer massive com- putational gains and parameter dimension reduction. An ingenious approach to Gaussian process construction in- volving the Matérn covariance function has been recently proposed by Li et al. [\[97\]](#page-20-26), who considered a product space involving the d-dimensional Euclidean space cross an ab-stract set that allows to index group labels.

## <sup>610</sup> **5.3 Directions, Shapes and Curves**

611 The Matérn model has an important role in the study 637 612 of directional processes, with Banerjee et al. [\[17\]](#page-18-27) formal- 638 ising the notions of directional *finite difference processes* and *directional derivative processes* with special empha-615 sis on the Matérn model. The Matérn model also has a role 641 616 in shape analysis, where Banerjee and Gelfand [\[15\]](#page-18-28) in-642 troduced *Bayesian wombling* to measure *spatial* gradients related to curves through 'wombling' boundaries, and ap-619 proach taken further in Halder et al. [\[70\]](#page-19-23). The smoothness properties of the Matérn model are ideally suited to such a framework. Modeling approaches to *temporal* gradients using the Matérn model have been proposed by Quick 623 et al. [\[132\]](#page-21-27). Related to these approaches, the smoothness parameter  $\nu$  of the Matérn model plays a central role in the recent paper by Halder et al. [\[70\]](#page-19-23), who analyse ran- dom surfaces in order to explain latent dependence within a response variable of interest.

This represents a short tour of *statistical* applications of the Matérn model, but its reach goes well beyond statistics, and we explore the importance of the Matérn model to related fields next.

# <span id="page-9-1"></span>**6. THE MATÉRN MODEL OUTSIDE STATISTICS**

This section explores the impact of the Matérn model <sup>633</sup> on numerical analysis and approximation theory (Section  $6.1$ ), machine learning (Section  $6.2$ ), and probability theory (Section  $6.3$ ).

# <span id="page-9-0"></span><sup>636</sup> **6.1 Numerical Analysis and Approximation Theory**

The problem considered here is to *reconstruct* a realvalued function f defined on a domain  $D \subset \mathbb{R}^d$  from given *data values*  $y_i = f(x_i)$  available at a set  $X_n =$  ${x_1, \ldots, x_n}$  of distinct *data locations*. In contrast to the statistical exposition in Section [3.1,](#page-3-0) from a numerical analysis standpoint these data are not assumed to be random in any way. Nevertheless, many of the mathematical expressions that we previously motivated from a statistical perspective appear also in the solution of this numerical task. The data vector  $Z_n$  is reinterpreted as  $\mathbf{Z}_n = (f(\mathbf{x}_1), \dots, f(\mathbf{x}_n))^{\top}$  and the task is to approximate the value  $f(x)$  of the unknown function f at an unsampled location  $x \in D \setminus X_n$ . A natural solution is a minimalnorm interpolant

$$
s_{f,X_n,K} = \underset{s \in \mathcal{H}(K)}{\arg \min} \|s\|_{\mathcal{H}(K)} \quad \text{s.t.} \quad \begin{array}{l} s(\boldsymbol{x}_i) = f(\boldsymbol{x}_i), \\ i = 1, \ldots, n, \end{array}
$$

which we recall was the third optimality property referred in Section [3.](#page-2-3) Thus, using again the *kernel matrix*  $\boldsymbol{R}_n = [K(\boldsymbol{x}_i, \boldsymbol{x}_j)]_{i,j=1}^n$ , the system  $\boldsymbol{R}_n \boldsymbol{b} = \boldsymbol{Z}_n$  is solved for a fixed coefficient vector  **that determines a linear** combination

$$
s_{f,X_n,K}(\boldsymbol{x}) = \sum_{i=1}^n b_i K(\boldsymbol{x}_i, \boldsymbol{x}), \qquad \boldsymbol{x} \in D,
$$

 $\sum_{i=1}^{\infty}$  in the span of the *translates*  $K(\mathbf{x}_i, \cdot)$ . This follows easily from the reproduction formula  $(3)$  and  $(2)$ . The above formula is identical to [\(12\)](#page-2-4) when setting  $x = x_0$ , and the resulting value  $s_{f,X_n,K}(x)$  is interpreted as a numerical approximation to  $f(x)$ . The log-likelihood function [\(13\)](#page-3-1) can equivalently be viewed as penalising the norm of the 643 interpolant, since  $||s_{f,X_n,K}||_{\mathcal{H}(K)}^2 = \mathbf{Z}_n^{\top} \mathbf{R}_n^{-1} \mathbf{Z}_n$ .

The fourth optimality principle in Section [3](#page-2-3) corresponds here to the fact that the norm of the error functional  $\epsilon_{\mathbf{x}} : f \mapsto f(\mathbf{x}) - s_{f,X_n,K}(\mathbf{x})$  in the dual space  $\mathcal{H}(K)^*$  of  $\mathcal{H}(K)$  is minimal under all linear reconstruction algorithms in  $\mathcal{H}(K)$  that use the same data  $\mathbf{Z}_n$ . The key tool is the *power function*  $P_{K,X_n}$ , defined for all  $x \in D$  by

$$
P_{K,X_n}(\boldsymbol{x})
$$
  
= sup { $f(\boldsymbol{x}) : f \in \mathcal{H}(K), f(X_n) = 0, ||f||_{\mathcal{H}(K)} \le 1$ }

It has the property  $P_{K,X_n}(\boldsymbol{x}) = ||\epsilon_{\boldsymbol{x}}||_{\mathcal{H}^*(K)}$  and leads to optimal error bounds of the form

$$
|f(\boldsymbol{x})-s_{f,X_n,K}(\boldsymbol{x})|\leq P_{K,X_n}(\boldsymbol{x})||f||_{\mathcal{H}_K}.
$$

644 for all  $x \in D$  and  $f \in \mathcal{H}(K)$ . It can be numerically cal-645 culated using the kernel matrix based on  $X_n \cup \{x\}$ , but 692 646 we omit the detail. Strikingly, the power function coin-<sup>647</sup> cides with the square root of the *kriging variance* [\[142\]](#page-21-28),  $_{648}$  giving the variance of the kriging error at x for given data  $_{695}$ 649 locations  $X_n$  and kernel K.

Analysis of the approximation error in this context thus  $_{697}$ reduces to analysis of the power function, and in turn analysis of the space  $\mathcal{H}(K)$ . From [\(4\)](#page-1-3) and [\(23\)](#page-5-4), the RKHS 699 generated by the Matérn kernel  $\mathcal{M}_{\nu,1}$  has the inner prod-  $_{700}$ uct

(27) 
$$
\langle f, g \rangle_{\mathcal{H}(\mathcal{M}_{\nu,1})} = \int_{\mathbb{R}^d} \frac{\hat{f}(\boldsymbol{\omega}) \overline{\hat{g}(\boldsymbol{\omega})}}{(1 + ||\boldsymbol{\omega}||^2)^{\nu + d/2}} d\boldsymbol{\omega}
$$

<sup>650</sup> up to constants, which we recognise as the inner product of the classical *Sobolev space*  $W_2^{\nu+d/2}$ <sup>651</sup> uct of the classical *Sobolev space*  $W_2^{\nu+d/2}(\mathbb{R}^d)$ . By the 652 Sobolev embedding theorem, the elements of this space 706 653 are well-defined continuous functions whenever  $\nu > 0$ . <sup>654</sup> This space is a canonical setting for mathematical anal-655 ysis of PDEs, a connection that we trailed in Section [4.4.](#page-7-0) 709 <sup>656</sup> Summarising, the use of Matérn kernels yields optimal 657 recovery techniques for functions in Sobolev spaces from 711 <sup>658</sup> given sampled data. Generalised recoveries using deriva-<sup>659</sup> tive data produce *meshless* numerical methods for solving <sup>660</sup> PDEs in Sobolev spaces, including the *symmetric colloca-*<sup>661</sup> *tion* method which uses derivative data for the PDE based  $\frac{662}{173}$  on Wu [\[173\]](#page-22-5), and shares similar Hilbert space optimality  $\frac{716}{216}$ 663 properties Schaback [\[138\]](#page-21-29). The use of the Matérn kernel 717 <sup>664</sup> is strongly motivated by the fact that PDE theory often 665 implies that solutions lie in Sobolev spaces. On the other 719 666 hand, there are also good arguments to replace Matérn 720 667 kernels by polyharmonics [\[139,](#page-21-30) [48\]](#page-19-24).

<sup>668</sup> Plenty of other results on deterministic recovery prob- $669$  lems using kernels can be found in Wendland [\[167\]](#page-21-1), while  $723$ 670 applications are in Schaback and Wendland [\[140\]](#page-21-31) and 724 671 MATLAB programs combined with the essential theory 725 <sup>672</sup> are in Fasshauer and McCourt [\[56\]](#page-19-25).

673 In numerical analysis and approximation theory, Matérn 727 <sub>674</sub> and other kernels are normally used for rather large val-675 ues of their smoothness parameter, because they seek to  $729$ 676 solve an interpolation rather than a regression task. Nar- 730  $677$  cowich et al.  $[115]$  proved that convergence rates then <sup>678</sup> depend on the minimum of the smoothness of the func-<sup>679</sup> tion f providing the data and the kernel; a *misspecified* 680 Matérn kernel, for which the smoothness parameter  $\nu$  is  $681$  taken to be too large relative to the smoothness of f, pro-<sup>682</sup> duces an error that converges at the same rate as we would 683 have achieved had  $\nu$  been correctly specified. On the other 684 hand, Tuo and Wang [\[160\]](#page-21-32) prove in the same setting that <sup>685</sup> the prediction error becomes more sensitive to the space-<sup>686</sup> filling property of the design points. In particular, optimal

<sup>687</sup> convergence rates require also that the *quasi-uniformity* of the experimental design is controlled.

<sup>689</sup> Of course, the use of kernels in numerical analysis and <sup>690</sup> approximation theory requires estimation of kernel pa-691 rameters. The quantity  $\sigma$  does not arise in the correlation matrix  $\mathbf{R}_n$ , but the scale parameter  $\alpha$  has a strong influence on the error of the interpolant. There is a vast literature on *scale estimation* that partially builds on statistical notions like ML (see references in Section 3). On <sup>696</sup> the other hand, specific alternatives to the Matérn model, such as the polyharmonic kernels of Section  $7.3$ , are able to bypass scale estimation due to the remarkable property that the interpolant is independent of the value of the scale parameter used. See Wendland  $[167]$  and Section [7.3.](#page-12-0)

## <span id="page-10-0"></span><sup>701</sup> **6.2 Machine Learning**

<sup>702</sup> Kernel methods are a major strand of machine learning <sup>703</sup> research, where kernels are routinely used to solve a vari-<sup>704</sup> ety of supervised and unsupervised learning tasks. Compared to the interpolatory setting of Section  $6.1$ , data in machine learning are usually observed with noise, necessitating either a likelihood or a loss function to be specified.

The Matérn model is often convenient for the analysis of kernel methods; for example, Tuo et al.  $[161]$  provide sufficient conditions for the rates of convergence of the Matérn kernel ridge regression to exceed the standard minimax rates under both the  $L_2$  norm and the norm of the RKHS. However, the presence of noise in the data can pose a substantial challenge to selection of smoothness parameters such as  $\nu$  in the Matérn model. Karvonen [\[80\]](#page-19-26) proves that the ML estimate of  $\nu$  cannot asymptotically *undersmooth* the truth under fixed domain asymptotics; that is, if the true regression function has a Sobolev smoothness  $\nu_0+d/2$ , then the smoothness parameter esti- $721$  mate cannot be asymptotically less than  $\nu_0 + d/2$ , but this in itself it not compelling motivation to use ML  $[81]$ . As a result of these additional challenges, standard practice is to keep the kernel general as far as possible when developing methodology, and as far as possible to learn a suitable <sup>726</sup> form for the kernel using the data and model selection criteria. However, recent machine learning methodology for non-Euclidean data hinges on the SPDE approach, and as a consequence the Matérn and related models are explicitly being used.

As the types of data that researchers seek to analyse become more heterogeneous and structured, there has been a demand for flexible Gaussian process models defined on such non-Euclidean domains as manifolds and discrete, graph-based domains. Under the framework of Gaussian processes, Borovitskiy et al. [\[34\]](#page-18-29) proposed to avoid numerical solution of the SPDE [\(24\)](#page-6-2) and instead to work with a finite-rank approximation to the Gaussian process model. Specifically, they consider the SPDE in [\(24\)](#page-6-2) appropriately adapted to a Riemannian manifold M, for which the corresponding Matérn model admits a series 779 expansion of the type

$$
\sum_{n=0}^{\infty} \left(\frac{2\nu}{\alpha^2} + \lambda_n\right)^{-\nu - d/2} f_n(\boldsymbol{x}) f_n(\boldsymbol{x}'), \qquad \boldsymbol{x}, \boldsymbol{x}' \in M
$$

<sup>731</sup> where  $\{\lambda_n\}_{n=0}^{\infty}$  and  $\{f_n\}_{n=0}^{\infty}$  are, respectively, the se- quences of eigenvalues and eigenfunctions from the Laplace–Beltrami operator  $-\Delta_M$ . The authors propose to first solve numerically for the leading eigenfunctions <sup>735</sup>  ${f_i}_{i=0}^n$  of the Laplace–Beltrami operator, and then work- ing with a finite-rank Gaussian process whose realisations <sup>737</sup> are linear combinations of the  $\{f_i\}_{i=0}^n$ . Though solving the eigenproblem may be harder than numerically solving the SPDE, the authors argue that caching of the eigen- functions can lead to a cost saving in settings where mul- tiple tasks are to be solved on the same manifold. Such an approach is ingeniously extended to undirected graphs by Borovitskiy et al.  $[33]$ , and has had a direct impact on Gaussian processes defined on neural networks [\[79\]](#page-19-28), pathwise conditioning of Gaussian processes [\[172\]](#page-22-6), sim- ulation intelligence in AI  $[91]$  and extension to kernel methods withing graphs cross time [\[116\]](#page-20-29). Other applica- tions include Thomson sampling in neural information systems [\[162\]](#page-21-34), Bayesian optimisation in robotics [\[78\]](#page-19-29), and Gaussian processes regression on metric spaces [\[89\]](#page-20-30).

## <span id="page-11-0"></span><sup>751</sup> **6.3 Probability Theory and Stochastic Processes**

 The Matérn model is well-studied from a probability theory and stochastic process viewpoint. From the per- spective of regularity, Scheuerer [\[141\]](#page-21-35) summarises the properties of Gaussian random fields with Matérn covari-756 ance functions; sample paths are  $k$ -times differentiable in  $\frac{1}{2}$  the mean-square sense if and only if  $\nu > k$ . Under the same condition, the sample paths have (local) Sobolev space exponent being identically equal to k. Further, a Gaussian random field with Matérn covariance has frac- tal dimension that is identically equal to  $\min(\nu, d)$ , for d being the dimension of the Euclidean space on which the random field is defined. For non-Gaussian random fields with Matérn covariance, continuity properties are studied by Kent [\[85\]](#page-19-30).

 Several other properties of the Matérn model have been investigated. Kelbert et al.  $[84]$  study fractional random  $812$  fields under the scenario of stochastic fractional heat equations under a Matérn model; see also Leonenko et al. [\[94\]](#page-20-31). Random fields defined on the unit ball embedded  $\mathbb{R}^d$ , with a covariance function that is the restriction of the Matérn model to a finite range, were studied in Leonenko et al. [\[93\]](#page-20-32). Tensor-valued random fields with  $818$  an equivalent class of Matérn covariance functions were 775 studied in Leonenko and Malyarenko [\[92\]](#page-20-33). Terdik [\[158\]](#page-21-36) 820 considers angular spectra for non-Gaussian random fields with Matérn covariance function. A recent contribution [\[159\]](#page-21-37) provides interesting connection between the Matérn

model and certain Laplacian ARMA representations of a class of stochastic processes. Lilly et al. [\[99\]](#page-20-34) show that the Matérn process is a damped version of fractional Brown- ian motion. Lim and Teo  $[100]$  study random fields with a generalised Matérn covariance obtained as the solution to the fractional stochastic differential equation with two fractional orders, enabling the authors to deduce the sample path properties of the associated random field. Space- time extensions of Matérn random fields through stochas-tic Helmholtz equations are provided by Angulo et al. [\[9\]](#page-18-31).

 According to N. Leonenko<sup>[1](#page-11-3)</sup>, a major contributor to this literature, the importance of Matérn model is based on the Duality theorem [\[71,](#page-19-32) Theorem 1] which provides an explicit relation between certain classes of characteristic functions of symmetric random vectors and their den- sity. Specifically, the spectral density associated with the Matérn model is by itself a covariance function, called the Cauchy or inverse multiquadric covariance function, that allows to parameterise the Hurst effect of the associated Gaussian random field.

This completes our tour across the scientific landscape 800 through the lens of the Matérn model. Our attention turns now to the future, and promising enhancements that can be made to the Matérn model.

## <span id="page-11-2"></span>**7. ENHANCEMENTS OF THE MATÉRN MODEL**

<sup>803</sup> This section described *enhancements* of the Matérn model; covariance functions that share (at least partially) 805 the local properties of the Matérn model while providing additional features and functionality. Here we first introduces the models one at a time, with critical commentary on their features deferred to Section [8.](#page-13-0)

## <span id="page-11-1"></span><sup>809</sup> **7.1 Models with Compact Support**

Compactly supported covariance models have a long history that can be traced back to Askey [\[11\]](#page-18-32), who proposed the kernel

(28) 
$$
\mathcal{A}_{\mu,\beta}(x) = \left(1 - \frac{x}{\beta}\right)_{+}^{\mu}, \qquad x \ge 0,
$$

810 with β and  $\mu$  being strictly positive, and where  $(x)_{+} =$  $\max(0, x)$  is the *truncated power*. It was shown in that work that  $\mathcal{A}_{\mu,\beta}$  belongs to  $\Phi_d$  for all  $\beta > 0$  if and only if  $\mu \ge (d+1)/2$ . Clearly, the mapping  $\mathbf{x} \mapsto \mathcal{A}_{\mu,\beta}(\|\mathbf{x}\|)$ is compactly supported over a ball with radius  $\beta$  embed- $_{815}$  ded in  $\mathbb{R}^d$ . As a result, covariance matrices contain exact zero entries whenever the associated states  $x_i$  and  $x_j$  satisfy  $||x_i - x_j|| \geq \beta$ ; the computational advantages of this sparsity are discussed further in Section [8.5.](#page-15-0)

Matheron's *montée* and *descente* [\[111\]](#page-20-10) approach was applied by Wendland  $[166]$  to the Askey functions, ob-821 taining compactly supported covariance functions with

<span id="page-11-3"></span><sup>1</sup> Personal Communication, January 2023.

 $\overline{2}$ 

822 higher-order smoothness that are truncated polynomials 823 as functions of  $||x||$ . This strategy was unable to generate <sup>824</sup> integer-order Sobolev spaces in even space dimensions, a  $825$  problem that was resolved in Schaback [\[137\]](#page-21-39) who identi-<sup>826</sup> fied the 'missing' Wendland functions. A unified view of  $827$  Wendland functions was provided by Gneiting [\[60\]](#page-19-18). Zas- $828$  tavnyi [\[176\]](#page-22-7) provided necessary and sufficient conditions 829 for a general class encompassing both ordinary and miss-830 ing Wendland functions. Buhmann [\[36\]](#page-18-33) provided a gener-831 alisation of Wendland functions, with sufficient paramet-832 ric conditions that allow the new class to belong to  $\Phi_d$ <sup>833</sup> for a given d. Those functions, termed *Buhmann* func- $834$  tions, were then studied by Zastavnyi [\[177\]](#page-22-8) and subse-835 quently by Zastavnyi and Porcu  $[178]$ , Porcu et al.  $[131]$ 836 and Faouzi et al. [\[54\]](#page-19-33). Alternative representations and 837 properties of the Wendland functions have been studied 862 838 by Hubbert [\[75\]](#page-19-34) and Chernih and Hubbert [\[38\]](#page-18-34). Exten-839 sions of the Wendland functions to multivariate  $[126, 44]$  $[126, 44]$  $[126, 44]$ ,  $_{864}$ 840 spatio-temporal [\[124\]](#page-20-37) and non-stationary processes [\[88\]](#page-20-38) <sup>841</sup> have also been developed.

842 A more technical discussion follows, in which we in-<sup>843</sup> troduce two further classes of correlation functions with 844 compact support, each of which will be the subject of dis-<sup>845</sup> cussion in Section [8.](#page-13-0)

> 1. The generalized Wendland  $(\mathcal{GW})$  family [\[59,](#page-19-35) [177\]](#page-22-8) contains correlation functions with compact support that, as in the Matérn model, admit a continuous parameterisation of smoothness of the underlying Gaussian random field. The  $\mathcal{GW}_{\kappa,\mu,\beta}$  model depends on parameters  $\kappa \geq 0$  and  $\mu, \beta > 0$  through the identity

(29) 
$$
\mathcal{GW}_{\kappa,\mu,\beta}(x) = \frac{\Gamma(\kappa)\Gamma(2\kappa+\mu+1)}{\Gamma(2\kappa)\Gamma(\kappa+\mu+1)2^{\mu+1}} \mathcal{A}_{\kappa+\mu,\beta^2}(x^2) \times {}_2F_1\left(\frac{\mu}{2},\frac{\mu+1}{2};\kappa+\mu+1;\mathcal{A}_{1,\beta^2}(x^2)\right),
$$

846 where  $\mu \ge (d+1)/2 + \kappa$  is needed for  $\mathcal{GW}_{\kappa,\mu,\beta}$ <sup>847</sup> to belong to the class  $\Phi_d$  and  ${}_2F_1(a, b, c, \cdot)$  is <sup>848</sup> the Gaussian hypergeometric function [\[2\]](#page-17-0). Sam-849 ple paths of the  $\mathcal{GW}_{\kappa,\mu,\beta}$  model are k times mean-<sup>850</sup> square differentiable, in any direction, if and only <sup>851</sup> if  $\kappa > k - 1/2$  [\[59\]](#page-19-35), so that  $\kappa$  plays the role of <sup>852</sup> the smoothness parameter in this model. When <sup>853</sup>  $\kappa = k \in \mathbb{N}, \mathcal{GW}_{k,\mu,\beta}$  factors into the product of <sup>854</sup> the Askey function  $A_{\mu+k,\beta}$  with a polynomial of  $\epsilon_{855}$  degree k. This model includes the Wendland func- $\epsilon_{\text{856}}$  tions ( $\kappa = k$ , a positive integer), as well as the  $\epsilon_{\text{879}}$ 857 missing Wendland functions ( $\kappa = k + 1/2$ ). The- $\frac{858}{23}$  orem 1(3) in Bevilacqua et al. [\[23\]](#page-18-18) implies that 859 the RKHS induced by  $\mathcal{GW}_{\kappa/2-(d+1)/4,\mu,\beta}$ , with 860  $\kappa \ge (d + 1/2)$ , is norm-equivalent to the Sobolev 861 space  $W_2^{\kappa}(\mathbb{R}^d)$ .

2. The Gauss hypergeometric  $(\mathcal{GH})$  family [\[51\]](#page-19-36) is defined as

<span id="page-12-2"></span>(30) 
$$
\mathcal{GH}_{\kappa,\delta,\gamma,\beta}(x) = \frac{\Gamma(\delta - d/2)\Gamma(\gamma - d/2)}{\Gamma(\delta - \kappa + \gamma - d/2)\Gamma(\kappa - d/2)} \times \mathcal{A}_{\delta - \kappa + \gamma - d/2 + 1, \beta^2}(x^2) \times 2F_1(\delta - \kappa; \gamma - \kappa; \delta - \kappa + \gamma - d/2; \mathcal{A}_{1,\beta^2}(x^2)).
$$

This model has four parameters and it belongs to the class  $\Phi_d$  for every positive  $\beta$  provided  $\kappa > d/2$ with

$$
2(\delta - \kappa)(\gamma - \kappa) \ge \kappa
$$
, and  $2(\delta + \gamma) \ge 6\kappa + 1$ .

Sample paths of the  $\mathcal{GH}_{\kappa,\delta,\gamma,\beta}$  model are  $\lceil k/2 \rceil$ 863 times mean-square differentiable, in any direction, if and only if  $\kappa > (k + d)/2$ . The parameter  $\kappa$  thus 865 also controls the smoothness of samples from this <sup>866</sup> model.

867 The importance of the GW and GH models is discussed 868 in Section [8.](#page-13-0)

#### <sup>869</sup> **7.2 Models with Polynomial Decay**

Correlation models with polynomial decay such as the generalized Cauchy [\[63\]](#page-19-37) or the Dagum models [\[20\]](#page-18-36) can be useful when modelling data with long-range dependence. However, in using these correlation models one loses control over the differentiability of the the sample paths, a key property of the Matérn model. Ma and Bhadra [\[107\]](#page-20-39) recently proposed a modification of the Matérn class that allows for polynomial decay, while maintaining the local properties of the conventional Matérn model. The correlation function associated to this model is given by

$$
(31)
$$

<span id="page-12-1"></span>
$$
\mathcal{CH}_{\nu,\eta,\beta}(x) = \frac{\Gamma(\nu+\eta)}{\Gamma(\nu)}\,\mathcal{U}\left(\eta,1-\nu,\nu\left(\frac{x}{\beta}\right)^2\right),\ x\geq 0,
$$

where  $U$  is the confluent hypergeometric function of the 871 second kind [\[2\]](#page-17-0). Here  $\nu > 0$  controls mean-square differ-872 entiability near the origin, as in the Matérn case, while  $873 \quad \eta > 0$  controls the heaviness of the tail. The construction  $(31)$  is based on a scale mixture of (a reparameterised version of) the Matérn model involving the inverse-gamma 876 distribution. Ma and Bhadra [\[107\]](#page-20-39) have shown that this 877 class is particularly useful for extrapolation problems where large distances are predominant.

# <span id="page-12-0"></span><sup>879</sup> **7.3 Polyharmonic Kernels**

Our catalogue of enhancements of the Matérn model finishes with *polyharmonic kernels*, defined as

<span id="page-12-3"></span>(32) 
$$
H_{\nu,d}(x) := \left\{ \begin{matrix} x^{2\nu - d} \log x & \text{for } 2\nu - d \in 2\mathbb{Z} \\ x^{2\nu - d} & \text{else} \end{matrix} \right\}
$$

<sup>880</sup> up to the sign  $(-1)^{\lfloor \nu - d/2 \rfloor + 1}$ . As a function of  $x = ||x||$ ,  $x \in \mathbb{R}^d$ , the Matérn kernel  $\mathcal{M}_{\nu-d/2,1}$  starts with even 882 powers of x followed by  $H_{\nu,d}$ , and in this sense the two <sup>883</sup> models are related. Up to a constant factor, the generalised Fourier transform of  $H_{\nu,d}(\Vert x \Vert)$  on  $\mathbb{R}^d$  is  $\Vert \omega \Vert^{-2\nu}$ , 885 and then a scale parameter is just another constant factor. <sup>886</sup> This makes kernel-based interpolation by polyharmonics  $887$  scale-independent. Compare with  $(23)$  to see the connec-888 tion to  $\mathcal{M}_{\nu-d/2,\alpha}$  in Fourier space. Stein [\[151\]](#page-21-41) provides <sup>889</sup> a formal connection between polyharmonic kernels, for <sup>890</sup> which the name *power law* covariance functions is also 891 used, and the Matérn model. Polyharmonic kernels are 892 *conditionally positive definite* of order  $|\nu - d/2| + 1$ ; 893 for a technical definition see Wendland  $[167]$ . Instead of <sup>894</sup> Hilbert Spaces, polyharmonic kernels generate *Beppo–* <sup>895</sup> *Levi spaces*, which share similarities to Sobolev spaces <sup>896</sup> modulo that an additional polynomial space has to be 897 added to enable prediction (Section [3\)](#page-2-3) and interpolation 898 (Section  $6.1$ ); see Wendland [\[167\]](#page-21-1). In general, polyhar-<sup>899</sup> monic kernels arise as covariances in *fractional* Gaussian <sup>900</sup> fields, including forms of Brownian motion [\[104,](#page-20-40) Theo-<sup>901</sup> rem 3.3]. 9  $\equiv \frac{1}{6}$   $\pm \frac{1}{6}$   $\pm$ 

<sup>902</sup> Next our attention turns to a critical discussion of <sup>903</sup> whether such enhancements to the Matérn model are <sup>904</sup> needed.

## **8. OTHER MODELS**

<span id="page-13-0"></span><sup>905</sup> This final section provides critical commentary on the <sup>906</sup> Matérn model and the enhanced versions of the model in-<sup>907</sup> troduced in Section [7.](#page-11-2)

#### <sup>908</sup> **8.1 Rigorous Generalisation of the Matérn Model**

<sup>909</sup> The Matérn model does not allow for compact support, <sup>910</sup> hole effects (oscillations between positive and negative 911 values) at large distances, or slowly decaying tails suit-<sup>912</sup> able for modeling long-range dependence. Most of the 913 enhancements in Section [7](#page-11-2) aim to resolve these kind of is-914 sues; here we describe how the GW, GH and CH models <sup>915</sup> can be viewed as rigorous generalisations of the Matérn <sup>916</sup> model.

Bevilacqua et al.  $[21]$  have shown that the Matérn  $_{921}$ model is a limit case of a rescaled version of the  $\mathcal{G}W$ model. In particular they have considered the model  $\mathcal{G}W$ defined as

$$
\mathcal{GW}_{\kappa,\mu,\beta}(x) = \mathcal{GW}_{\kappa,\mu,\beta\left(\frac{\Gamma(\mu+2\kappa+1)}{\Gamma(\mu)}\right)^{\frac{1}{1+2\kappa}}}(x), \qquad x \ge 0,
$$

and proved that

$$
\lim_{\mu\to\infty}\widetilde{\mathcal{G}\mathcal{W}}_{\kappa,\mu,\beta}(x)=\mathcal{M}_{\kappa+1/2,\beta}(x),\quad \kappa\geq 0,
$$

917 with uniform convergence over the set  $x \in (0, \infty)$ . Figure 9[1](#page-13-1)8 1 (first row) depicts the convergence result for  $\nu = 0, 1, 2$ 



<span id="page-13-1"></span>FIG 1*. First row: the*  $\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}(x)$  *model with*  $\mu = 3, 5, 15$  *and*  $\mu \rightarrow \infty$  *(the Matérn model*  $\mathcal{M}_{\kappa+1/2,\beta}(x)$ *)* with  $\beta = 0.15, 0.12, 0.1$ and  $\kappa = 0, 1, 2$  *(from left to right) respectively. Second row: the*  $\mathcal{CH}_{\nu,\eta,2\sqrt{\nu(\eta+1)}\beta}(x)$  *model with*  $\eta=3,5,15$  *and*  $\eta\to\infty$ *(the Matérn model*  $\mathcal{M}_{\nu,\beta}(x)$ *) with*  $\beta = 0.15, 0.12, 0.1$  *and*  $\nu =$ 0.5, 1.5, 2.5 *(from left to right) respectively.*

The parameter  $\mu$  thus allows for switching from compactly supported to globally supported models, and can either be fixed to ensure sparse correlation matrices, or can be estimated based on the dataset. However, this equivalence applies only to smoothness parameters greater than or equal to  $1/2$  in the Matérn model, so the full range of the smoothness parameter is not covered. This is unfortunate, since the fractal dimension [a widely used measure of roughness of the sample paths for time series and spatial data; [64\]](#page-19-38) is fully parameterised using the Matérn model when the smoothness parameter lies between 0 and 1. As a consequence, the  $\mathcal{GW}$  (or  $\mathcal{GW}$ ) model cannot fully parameterise the fractal dimension of the random field. This kind of issue can be solved with the  $\mathcal{GH}$  model, which includes the  $\mathcal{GW}$  model as a special case  $[51]$ :

$$
\mathcal{GH}_{\frac{d+1}{2}+\nu,\frac{d+\mu+1}{2}+\nu,\frac{d+\mu}{2}+1+\nu,\beta}(x)=\mathcal{GW}_{\nu,\mu,\beta}(x)
$$

920 Letting  $\beta$ , δ and  $\gamma$  tend to infinity in such a way that <sup>920</sup> Letting *p*, *o* and *γ* tend to minity in such a way that  $\beta/\sqrt{4\delta\gamma}$  tends to  $\alpha > 0$ , the *GH* model [\(30\)](#page-12-2) converges 922 uniformly to the Matérn model  $\mathcal{M}_{\kappa-d/2,\alpha}(x)$ , and in this case the *full range* of the smoothness parameter of the <sup>924</sup> Matérn model is covered.

The Matérn model also arises as a special limit case of the  $\mathcal{CH}$  model. Specifically, Ma and Bhadra [\[107\]](#page-20-39) show that

$$
\lim_{\eta \to \infty} \mathcal{CH}_{\nu, \eta, 2\sqrt{\nu(\eta+1)}\beta}(x) = \mathcal{M}_{\nu, \beta}(x),
$$

925 with convergence being uniform on any compact set. Fig-926 ure [1](#page-13-1) (second row) depicts the convergence result for  $\nu = 0.5, 1.5, 2.5$  (from left to right).

<sup>928</sup> The *turning band* operator of Matheron [\[110\]](#page-20-9) can be 929 applied to a correlation function to create hole effects <sup>930</sup> while retaining positive definiteness of the kernel. An ar-931 gument in Schoenberg proves that, for an isotropic corre-932 lation in  $\mathbb{R}^d$ , the correlation values cannot be smaller than 933  $-1/d$  [\[144\]](#page-21-42). Since the Matérn model is a valid model for  $934$  all d, this implies that the application of turning bands to 935 the Matérn model will not provide any hole effect. On the 936 other hand, the GW and GH models allow for such an <sup>937</sup> effect.

## <sup>938</sup> **8.2 Estimation of Enhanced Models**

939 ML estimation for the Matérn model are well-understood; <sup>940</sup> here we discuss the extent to which similar results can be 941 obtained for enhancements of the Matérn model.

 In the context of increasing domain asymptotics, pa-943 rameters of the  $\mathcal{GW}$  and  $\mathcal{CH}$  models can be estimated consistently using ML and the associated asymptotic dis-tribution is known; see Section [3.1.1.](#page-3-3)

In the context of fixed domain asymptotics, similar to the classical Matérn model, the parameters of the these enhanced models cannot be consistently estimated. For instance, Bevilacqua et al. [\[23\]](#page-18-18) show that the microergodic parameter of the covariance model  $\sigma^2 \mathcal{GW}_{\kappa,\mu,\beta}$ , assuming  $\kappa$  and  $\mu$  known, is given by micro $\varsigma_W$  =  $\sigma^2/\beta^{2\kappa+1}$ . In addition they prove that for a zero mean Gaussian field defined on a bounded infinite set  $D \subset \mathbb{R}^d$  $(d = 1, 2, 3)$ , with covariance model  $\sigma_0^2 \mathcal{GW}_{\kappa,\mu,\beta_0}$ , the ML estimator  $\hat{\sigma}_n^2 / \hat{\beta}_n^{2\kappa+1}$  of the microergodic parameter is strongly consistent, i.e.,

$$
\hat{\sigma}_n^2/\hat{\beta}_n^{2\kappa+1} \xrightarrow{a.s.} \sigma_0^2/\beta_0^{2\kappa+1}.
$$

Additionally, for  $\mu > (d+1)/2 + \kappa + 3$ , its asymptotic distribution is given by

$$
\sqrt{n}(\hat{\sigma}_n^2/\hat{\beta}_n^{2\kappa+1} - \sigma_0^2/\beta_0^{2\kappa+1}) \stackrel{d}{\longrightarrow} \mathcal{N}(0, 2(\sigma_0^2/\beta_0^{2\kappa+1})^2).
$$

946 Analogous for the  $\mathcal{GH}$  model proposed are not available <sup>947</sup> at present.

Similarly, Ma and Bhadra [\[107\]](#page-20-39) show that the microergodic parameter of the covariance model  $\sigma^2 C \mathcal{H}_{\nu, \eta, \beta}$ , assuming  $\nu$  known, is given by

$$
\text{micro}_{\mathcal{CH}} = (\sigma^2 \Gamma(\nu + \eta)) / (\beta^{2\nu} \Gamma(\eta)).
$$

In addition they prove that for a zero mean Gaussian field  $_{963}$ defined on a bounded infinite set  $D \subset \mathbb{R}^d$   $(d = 1, 2, 3)$ , with covariance model  $\sigma_0^2 C \mathcal{H}_{\nu, \eta_0, \beta_0}$ , the ML estimator  $(\hat{\sigma}_n^2/\hat{\beta}_n^{2\nu})(\Gamma(\nu + \hat{\eta}_n)/\Gamma(\hat{\eta}_n))$  of the microergodic parameter is strongly consistent, i.e.,

$$
\frac{\hat{\sigma}_n^2(\Gamma(\nu + \hat{\eta}_n)}{\hat{\beta}_n^{2\nu}\Gamma(\hat{\eta}_n)} \xrightarrow{a.s.} \frac{\sigma_0^2\Gamma(\nu + \eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)}
$$

and, if  $\eta_0 > d/2$ , its asymptotic distribution is given by

$$
\frac{\hat{\sigma}_n^2(\Gamma(\nu + \hat{\eta}_n))}{\hat{\beta}_n^{2\nu}\Gamma(\hat{\eta}_n)} - \frac{\sigma_0^2\Gamma(\nu + \eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)} \longrightarrow \mathcal{N}\left(0, 2\left(\frac{\sigma_0^2\Gamma(\nu + \eta_0)}{\beta_0^{2\nu}\Gamma(\eta_0)}\right)^2\right).
$$

948 These results broadly support the use of ML plug-in esti-949 mates for these enhanced versions of the Matérn model; <sup>950</sup> the issue of predictive performance is discussed next.

#### <sup>951</sup> **8.3 Prediction with Enhanced Models**

If two Gaussian measures are equivalent then the associated predictions and mean squared errors are asymptotically identical (c.f. Section [3.2\)](#page-4-0). To this end, recent results have sought to establish equivalence between Gaussian measures for the Matérn model and enhancements of the Matérn model. Bevilacqua et al. [\[23\]](#page-18-18) consider the  $\sigma_1^2 \mathcal{GW}_{\kappa,\mu,\beta}$  model and show that for given  $\sigma_1 \geq 0$ ,  $\nu \ge 1/2$ , and  $\kappa \ge 0$ , if  $\nu = \kappa + 1/2$ ,  $\mu > d + \kappa + 1/2$  and

$$
(33) \qquad \sigma_0^2 \alpha^{-2\nu} = \left(\frac{\Gamma(2\kappa + \mu + 1)}{\Gamma(\mu)}\right) \sigma_1^2 \beta^{-(1+2\kappa)},
$$

then  $P(\sigma_0^2 M_{\nu,\alpha})$  is equivalent to  $P(\sigma_1^2 \mathcal{GW}_{\kappa,\mu,\beta})$ , for  $d=$ 1, 2, 3, on the paths of  $Z(x)$  for  $x \in D \subset \mathbb{R}^d$ . Thus predictions made using the  $\mathcal{GW}$  model with compact support are asymptotically identical to those made using the Matérn model. Likewise, Ma and Bhadra [\[107\]](#page-20-39) show that for a given  $\eta \ge d/2$  and  $\nu \ge 0$ , if

(34) 
$$
\sigma_0^2 \alpha^{-2\nu} = \left(\frac{\Gamma(\nu + \eta)}{\Gamma(\eta)}\right) \sigma_1^2 \left(\frac{\beta^2}{2}\right)^{-\nu},
$$

<sup>952</sup> then  $P(\sigma_0^2 M_{\nu,\alpha})$  is equivalent to  $P(\sigma_1^2 C H_{\nu,\eta,\beta})$ , for  $d =$ 953 1, 2, 3, on the paths of  $Z(x)$  for  $x \in D \subset \mathbb{R}^d$ . Thus pre-954 dictions made using the  $\mathcal{GW}$  model with polynomial tail <sup>955</sup> decay are asymptotically identical to those made using the <sup>956</sup> Matérn model.

 $_{957}$  If interest is in the predictor [\(12\)](#page-2-4), but not the predictive <sup>958</sup> uncertainty resulting from the associated Gaussian ran-<sup>959</sup> dom field, then it is interesting to note that the stationar-<sup>960</sup> ity assumption of the Matérn model may not be needed. 961 Stein et al. [\[154\]](#page-21-43) showed that, under suitable paramet-962 ric conditions, one can consider  $\alpha = 0$  in the Matérn model, and this is equivalent to prediction using the polyharmonic kernels  $H_{\nu,d}$  in [\(32\)](#page-12-3). Theorem 1 in that work shows that if  $d \leq 3$  and the parameter  $\nu$  satisfies condition (2) therein (or  $d = 1$ ), then *it is impossible to distin-*967 *guish*  $\alpha > 0$  *from*  $\alpha = 0$  *on a bounded domain.* The above <sup>968</sup> observation reflects the fact that prediction using polyhar-969 monic kernels, like in Section [6.1,](#page-9-0) is scale-independent. 970 This follows from homogeneity of the Fourier transform 971 and eliminates the need for scale estimation in this con-<sup>972</sup> text.

# <sup>973</sup> **8.4 Screening with Enhanced Models**

974 The screening effect extends also to enhanced versions<sup>1027</sup> 975 of the Matérn model. For regular schemes, Theorem 1 in<sup>1028</sup> 976 Porcu et al. [\[130\]](#page-21-11) shows that the  $\mathcal{GW}$  model allows for an <sup>1029</sup> 977 asymptotic screening effect when  $\mu > (d+1)/2 + \kappa$ . This 978 condition is not restrictive, since  $\mu \ge (\dot{d} + 1)/2 + \nu$  is al-<sup>1031</sup> 979 ready required for  $\mathcal{GW}_{\kappa,\mu,\beta}$  to belong to the class  $\Phi_d$ . For <sup>1032</sup> 980 irregular schemes the situations is more complicated. For <sup>1033</sup> 981 example, for non-differentiable fields in  $d = 1$ , Theorem <sup>1034</sup> 982 1 in Stein [\[152\]](#page-21-9) in concert with Theorem 1 in Porcu et al. <sup>1035</sup> 983 [\[130\]](#page-21-11) explains that the Askey model  $\mathcal{GW}_{0,\mu,\beta}$  allows for <sup>1036</sup> 984 a screening effect provided  $\mu > 1$ . For  $d = 2$ , Theorem <sup>1037</sup> 985 2 in Stein [\[152\]](#page-21-9) implies that the Askey model allows for <sup>1038</sup> 986 screening provided that  $\mu > 3/2$ . The GW model satisfies <sup>1039</sup> 987 Stein's condition in  $(1.3)$  of Porcu et al.  $[130]$ , which in <sup>1040</sup> 988 turn allows the Stein hypothesis  $(22)$  to be verified.

989 The numerical experiments in Porcu et al. [\[130\]](#page-21-11) sug-<sup>1042</sup> 990 gest that the screening effect is even stronger under en-<sup>1043</sup> 991 hanced models with compact support, compared to the <sup>1044</sup> 992 standard Matérn model. This can deliver computational 1045 <sup>993</sup> advantages, which we discuss next.

## <span id="page-15-0"></span><sup>994</sup> **8.5 Scalable Computation**

995 Scalable computation generally refers to the computa-<sup>996</sup> tional complexity associated with the optimal predictor  $_{997}$  [\(12\)](#page-2-4) and/or of the likelihood function [\(13\)](#page-3-1) when increas-998 ing *n* the number of data.

999 The eternal fight between statistical accuracy and com- putational scalability has produced methods that attempt to deal with this notorious trade-off. The discussion that follows focuses specifically on this trade-off in the con- text of the Matérn model or its enhancements, in par- ticular compactly supported models. General approaches, such as those based on predictive processes [\[16\]](#page-18-37) and those based on fixed-rank kriging [\[43\]](#page-18-38), will not be discussed; the interested reader is referred to the review of Sun et al. <sup>1008</sup> [\[155\]](#page-21-20).

1009 The computational complexity associated with the <sup>1062</sup> 1010 Matérn model is broadly governed by the number of data<sup>1063</sup> 1011 (n) and partially by the input space dimension  $(d)$ , the <sup>1064</sup> 1012 dimension,  $p$  of the (scalar or vector) random field. In the <sup>1065</sup> <sup>1013</sup> case of scalar-valued random fields, we have  $p = 1$ .

<sup>1014</sup> These challenges will be considered in turn.

1015 The flexibility of some enhanced models is lost in the <sup>1068</sup> 1016 case of large domains; the condition  $\mu \ge (d+1)/2 + \kappa^{1069}$ 1017 in the  $\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}$  model forces the parameter  $\mu$  to go to in-1018 finity with d, which in turn forces  $\widetilde{\mathcal{GW}}_{\kappa,\mu,\beta}$  to approach  $\frac{1}{1072}$ <br>1019  $\mathcal{M}_{\mu,\alpha}$ . From this point of view the class  $\mathcal{GH}_{\kappa,\delta,\alpha}$  seems  $\mathcal{M}_{\nu,\alpha}$ . From this point of view the class  $\mathcal{GH}_{\kappa,\delta,\gamma,\beta}$  seems  $_{1073}$ 1020 more promising to use for large  $d$ . An additional remark  $_{1074}$ 1021 is that, for  $d \geq 5$ , all Gaussian measures with Matérn co- $\frac{1000}{1075}$ <sup>1022</sup> variance functions are orthogonal [\[7\]](#page-18-2). This has philosoph-<sup>1023</sup> ical consequences for Gaussian process regression when <sup>1024</sup> the Matérn model is viewed as a *prior* distribution en-<sup>1025</sup> coding *a priori* belief, since a small change to the kernel

<sup>1026</sup> parameters results in the entire support of the prior being changed.

In the case of a large number of variables  $p$  in a multivariate the model, a large number of parameters needs to be estimated.

The multivariate Matérn model suffers from the fact, not only does the number of parameters increase polynomially with  $p$ , but the conditions for validity of the model imply severe restrictions on the collocated correlation coefficient  $\rho_{ij}$  in [\(26\)](#page-8-1). Emery et al. [\[52\]](#page-19-21) show that such restrictions become extremely severe already with  $p = 3$ . Similar comments apply to other multivariate covariance functions, including the multivariate  $\mathcal{GW}$  model in Daley et al.  $[44]$ .

Finally we consider the case where the number  $n$  of <sup>1041</sup> data is large, entailing a  $O(n^3)$  computational and  $O(n^2)$ storage cost associated with the predictor  $(12)$  or the likelihood function( $13$ ). Several approaches have been proposed to reduce these costs, many of which take advantage of the (approximate) sparsity of the covariance  $(\Sigma_n)$ 1046 or precision  $(\mathbf{\Sigma}_n^{-1})$ , or its Cholesky factor  $(\text{ch}(\mathbf{\Sigma}_n^{-1}))$ :

- $\bullet$  Sparsity in the covariance matrix  $\Sigma_n$  can be di-<sup>1048</sup> rectly exploited by using compactly supported <sup>1049</sup> models such as the  $\mathcal{GW}_{\kappa,\mu,\beta}$  or the  $\mathcal{GH}_{\kappa,\delta,\gamma,\beta}$  fam-<br><sup>1050</sup> ilies. Such approaches can be useful when the (esilies. Such approaches can be useful when the (es-<sup>1051</sup> timated) compact support is relatively small with <sup>1052</sup> respect to the spatial extent of the sampling region, <sup>1053</sup> so that approximations are extremely sparse; see <sup>1054</sup> below for an empirical investigation of this point.
- <sup>1055</sup> The precision matrix  $\Sigma_n^{-1}$  associated with the <sup>1056</sup> Matérn model is in general non-sparse (except for the case  $d = 1$  and  $\nu = 0.5$ ) but it turns out that the <sup>1058</sup> matrix values are in general relatively close to 0, i.e.  $\Sigma_n^{-1}$  is *quasi-sparse*. As a consequence, approxi-1060 mating  $\Sigma_n^{-1}$  with a sparse matrix can be a good <sup>1061</sup> strategy. A notable instance of this approach is the SPDE approach from Section  $4.2$ . This approach can be also motivated from results in numerical linear algebra, which demonstrate that if the elements of a matrix show a property of decay, then the ele-<sup>1066</sup> ments of its inverse also show a similar (and faster)  $1067$  behavior [\[19\]](#page-18-39).
- Vecchia's approximation  $[164]$  and its extensions [e.g.  $46, 66, 82, 45$  $46, 66, 82, 45$  $46, 66, 82, 45$  $46, 66, 82, 45$  $46, 66, 82, 45$  $46, 66, 82, 45$ ] imply a sparse approximation <sup>1070</sup> of ch( $\Sigma_n^{-1}$ ) and are often applied to the Matérn model, although they can be applied to any covariance model. One potential limitation of these method is that they depend on an ordering of the variables and the choice of conditioning sets which determines the Cholesky sparsity pattern [see [66\]](#page-19-12).

It is instructive to numerically investigate the sparseness of matrices associated with enhancements of the Matérn model, and for this we focus on the  $\mathcal{GW}_{\kappa,\mu,\beta}$  model,





 $#∞ ≌ ‰ ন ৪$ 

<span id="page-16-0"></span>*Sparsity (percentage of zero values in the upper triangular part) of the covariance matrix* Σn*, and quasi-sparsity (defined in the main text)* in the precision matrix  $(\boldsymbol{\Sigma}_n^{-1})$  and its Cholesky factor  $(ch(\boldsymbol{\Sigma}_n^{-1}))$  for *the* GW<sub>K,µ,β</sub> *model. The case* GW<sub>K,∞,β</sub> corresponds to the Matérn *model*  $\mathcal{M}_{\nu+1/2,\beta}$ *. The*  $\beta$  *parameters are chosen so that the practical range of the Matérn model is equal to* 0.15*.*

$\kappa = 2$	$\mathrm{ch}(\boldsymbol{\Sigma}_{n}^{-1})$	4489						$0.90$ $0.91$ $0.914$ $0.946$ $0.9$	
		1156				213 21435.9			25.9
	$\Sigma_n^{-1}$					$2.73$ $25.8$	23.5	10.9	18.2
		4489 1156 4489			$0.71\,$	$13.7\,$	25.5	$26.0\,$	$26.0\,$
	$\Sigma_n$		71.4	66.4	29.6	1.29			
		1156	72.6	67.2	30.7	1.72			
	$\begin{array}{c c} & \mu & \end{array}$		0.35	0.39	$0.67\,$	1.21	2.29	8.24	8
			3.5			$\overline{16}$	32	120	$\,8\,$
$\kappa = 1$	$ch(\Sigma_n^{-1})$		$\overline{a}$					8.3 0.21 8.4 9.8 0.21 8.4 9.8	
			$\ddot{\phantom{0}}$	1.12		$4.49$ $17.1$ $29.2$ $32.3$			34.3
		156   4489   1156   4489   1156   4489				$-23.5$ $-14.3.5$ $+5.3$		59.8	$61.3\,$
	$\Sigma_n^{-1}$							$\begin{array}{c} 5.91 \\ 2.1.2 \\ 3.7.2 \\ 42.6 \end{array}$	
			80.2	63.7	19.7				
	$\Sigma_n$		$80.5$	65.0	<b>20.7</b>				
	$\begin{array}{c c} \hline \mu & C \end{array}$					$0.28$ $0.42$ $0.5$ $0.43$ $0.2$ $1.42$ $0.2$			8
			$\frac{15}{25}$		∞	$\frac{6}{1}$	32	120	8
$\kappa=0$	$ch(\Sigma_n^{-1})$				2.15 2.15 21.3		24.3	21.9	24.0
		1156   4489		0.58		$1.84$ $4.98$ $12.7$		8.33	9.85
	$\Sigma_n^-$	4489	$\overline{0}$			0.0 0.0.0.4.4 0.0.4.4.4			47.9
								1110801 11521123	
	$\mathbf{z}_n$	156   4489   1156			$\begin{array}{c} 89.1 \\ 47.3 \\ 1.21 \end{array}$				
			$\begin{array}{c c c}\n 0.20 & 90.0 & \circ. \\  -48.7 & 47.2\n\end{array}$		1.43				
						$\begin{array}{c} 1.07 \\ 2.13 \\ 4.27 \\ 16.0 \end{array}$			$\,$ $\,$
			1.5					$\overline{20}$	

<span id="page-16-1"></span>TABLE 2 *As in Table [1,](#page-16-0) but with* β *chosen such that the practical range of the Matérn model is equal to 0.4.*

which allows us to switch from a model with compact support of radius

$$
C = \beta \left( \frac{\Gamma(\mu + 2\kappa + 1)}{\Gamma(\mu)} \right)^{\frac{1}{1+2\kappa}}
$$

1076 to the Matérn model by increasing the  $\mu$  parameter. In <sup>1077</sup> our experiment, the sparseness of  $\Sigma_n$  and the *quasi-*<sup>1078</sup> *sparseness* of  $\Sigma_n^{-1}$  and ch $(\Sigma_n^{-1})$  are reported, the latter 1079 being defined as the percentage of values in the upper tri-<sup>1132</sup> 1080 angular matrix with absolute value lower than an arbitrary<sup>1133</sup> 1081 small constant  $\epsilon$ , and in our example we set  $\epsilon = 1.e - 8$ . 1082 The empirical assessment considers  $n = 1,156$  and <sup>1135</sup> 1083  $n = 4,489$  location sites over  $[0,1]^2$ , where the points are 1084 equally spaced by 0.03 and 0.015 respectively in a reg- $1137$ 1085 ular grid. For  $\nu = 0, 1, 2$ , we set  $\beta$  such that the prac-<sup>1138</sup> 1086 tical range of the Matérn model is equal to  $0.15 \bar{(\beta)} = 1139$ 1087 0.050, 0.0316, 0.0253 respectively), and consider increas-<sup>1140</sup> <sup>1088</sup> ing  $\mu = 1.5 + \kappa, 4, 8, 16, 32, 120, \infty$  (with  $\widetilde{\mathcal{GW}}_{\kappa,\infty,\beta}$  being <sup>1141</sup><br><sup>1089</sup> the Matérn model  $\mathcal{M}_{\kappa+1/2,\beta}$ ). the Matérn model  $\mathcal{M}_{\kappa+1/2,\beta}$ ).

1090 The results are reported in Table [1.](#page-16-0) For the low val-<sup>1143</sup> 1091 ues  $\mu = 1.5, 2.5, 3.5$  and  $\nu = 0, 1, 2$ , the covariance ma-<sup>1144</sup> 1092 trix is highly sparse, while the sparseness decreases when  $1145$ 1093 increasing  $\mu$ , as expected. There is a clear trade-off be-<sup>1146</sup> tween the sparseness of  $\Sigma_n$  and quasi-sparseness of  $\Sigma_n^{-1}$ 1094 1095 and ch $(\Sigma_n^{-1})$  for each  $\nu = 0, 1, 2$ . However, when increas-1096 ing  $\mu$ , that is when  $\Sigma_n$  approaches the Matérn covariance  $\lim_{n \to \infty}$  matrix, then  $\Sigma_n^{-1}$  or ch $(\Sigma_n^{-1})$  tends to be highly quasi-<sup>1098</sup> sparse.

1099 We replicate the same experiment but with a practical <sup>1152</sup> 1100 range of the Matérn model equal to 0.4. This leads to <sup>1153</sup> 1101  $\beta = 0.133, 0.084, 0.067$  for  $\nu = 0, 1, 2$  respectively. The 1154 <sup>1102</sup> results are reported in Table [2.](#page-16-1) The conclusions are the <sup>1103</sup> same of the previous setting but in this case, we have 1104 lower levels of sparseness for  $\Sigma_n$  and of quasi-sparseness 1105 for  $\Sigma_n^{-1}$  and  $\text{ch}(\Sigma_n^{-1})$ .

<sup>1106</sup> These numerical experiments highlight a clear trade-off 1107 between the (quasi-)sparseness of  $\sum_{n}^{-1}$  (or ch( $\sum_{n}^{-1}$ )) and <sup>1108</sup>  $\Sigma_n$  when increasing  $\mu$  for fixed  $\beta$  and  $\nu$  i.e. when switch-<sup>1109</sup> ing from a compactly supported to a globally supported 1110 Matérn model. In particular, when  $\mu \to \infty$  (the Matérn 1111 model), then  $\Sigma_n^{-1}$  is highly quasi-sparse and  $\Sigma_n$  is dense. not model), then  $\mathbb{Z}_n$  is highly quasi-sparse and  $\mathbb{Z}_n$  is dense.<br>1112 In contrast, when  $\mu$  is small then  $\mathbb{Z}_n^{-1}$  is not quasi-sparse 1113 yet  $\Sigma_n$  is highly sparse. This seems to suggest that sparse  $\frac{n}{1163}$ 1114 precision matrix approximation should work reasonably <sup>1164</sup> 1115 well for the Matérn model, but could be problematic when <sup>1165</sup> <sup>1116</sup> handling data exhibiting short compactly supported de-1117 pendence. In this case a better approach should be to ex-1118 ploit the sparsity of  $\Sigma_n$ , as enabled by enhanced versions <sup>1119</sup> of the Matérn model.

1120 As a final comment, the evaluation of the Matérn model 1168 1121 and its enhancements requires the computation of some<sup>1169</sup>  $1122$  special functions such as the modified Bessel function of  $\frac{1170}{1171}$  $1123$  the second kind, the Gaussian hypergeometric function  $\frac{1172}{1172}$  $_{1124}$  and the confluent hypergeometric function of the second  $_{1173}$ 1125 kind that can be found in different libraries such as the 1174  $1126$  GNU scientific library  $[65]$  and the most important sta- $1175$ <sup>1127</sup> tistical softwares including R, MATLAB and Python. For 1128 instance, the R package GeoModels  $[25]$  implements the  $\frac{1177}{1178}$ 1129 computation of the Matérn model and its enhancements  $\frac{1179}{1179}$ 1130 for  $d = 1, 2$ .

## **9. CONCLUSION**

The impact of the Matérn model since its conception has been substantial, and the model continues to be widely used, across a broad range of scientific disciplines and be-<sup>1134</sup> yond. While the original motivation for the Matérn model came from its flexibility in context of spatial interpolation, there is now also a rich literature of alternative and enhanced versions of the model. In particular, the SPDE and related approaches enable one to define analogues of the Matérn model on quite general domains, admitting sparse approximations to precision matrices, while recent advances in enhanced models with compact support can facilitate scalable computation through sparse approximation of covariance matrices, and are well-suited to processes with short-scale dependence. The theoretical and empirical properties of these enhanced models have been recently and actively studied. On the other hand, there remain open theoretical issues of practical importance, such as parameter estimation at finite sample sizes, and the impact of parameter estimation on the performance of the associated predictions.

Our current understanding of the Matérn model has emerged as the result of engagement between scientists and practitioners from different disciplines, and our hope is that this multi-disciplinarity perspective will shine further light onto the Matérn model.

#### **ACKNOWLEDGMENTS**

We thank the Associate Editor and three anonymous Reviewers for their thorough reading and criticisms that <sup>1158</sup> allowed for an improved version of the manuscript. We are very grateful to Toni Karvonen for pointing out an important technicality about Sobolev spaces associ-<sup>1161</sup> ated with the Matérn kernel. Moreno Bevilacqua acknowledges financial support from grant FONDECYT <sup>1163</sup> 1240308 and ANID/PIA/ANILLOS ACT210096 and ANID project Data Observatory Foundation DO210001 from the Chilean government and project MATH-AMSUD <sup>1166</sup> 22-MATH-06 (AMSUD220041).

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